

*Credit : Max Alexander
Airbus Defence and Space*



Identifying heterogeneity sizes in the subsurface with WISDOM

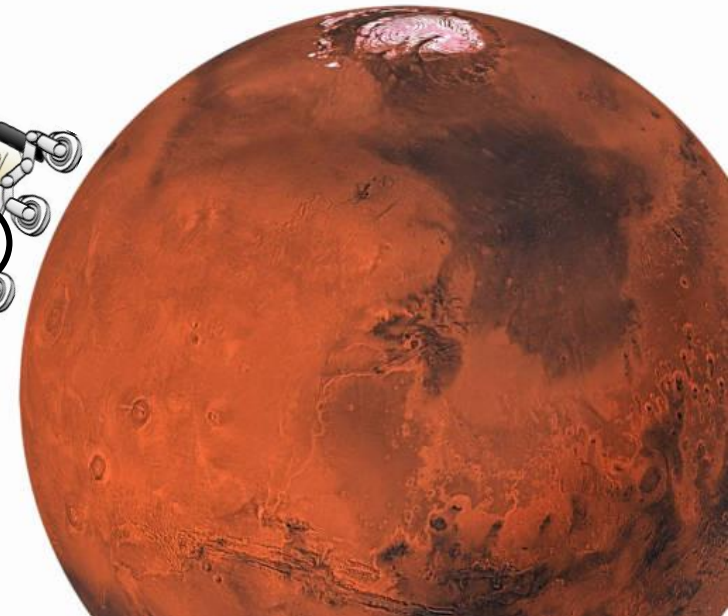
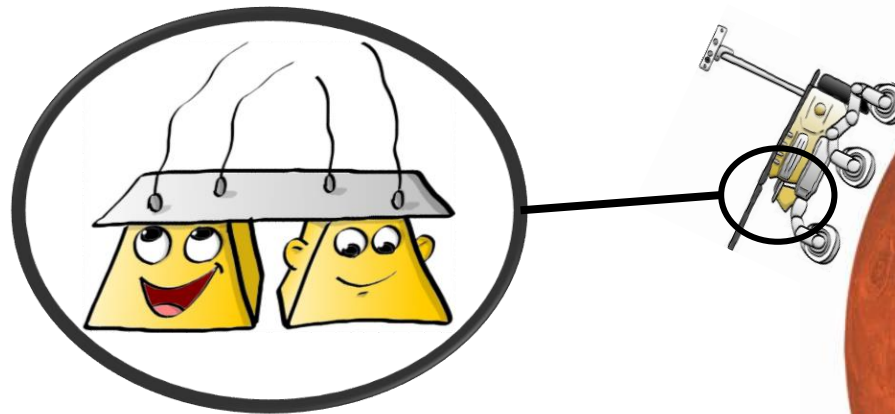
The Ground Penetrating Radar of the
ExoMars Martian mission

Author : Emile Brighi

Contact : emile.brighi@latmos.ipsl.fr

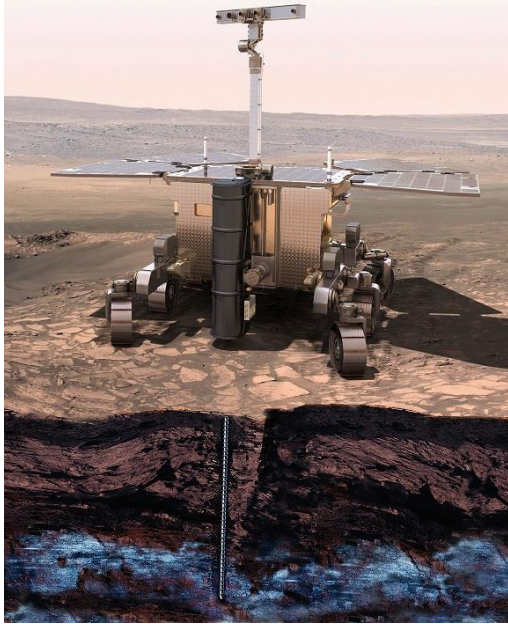
Co-authors :

- Valérie Ciarletti
- Alice Le Gall
- Nicolas Oudart
- Yann Herve



THE ROSALIND FRANKLIN / EXOMARS MISSION

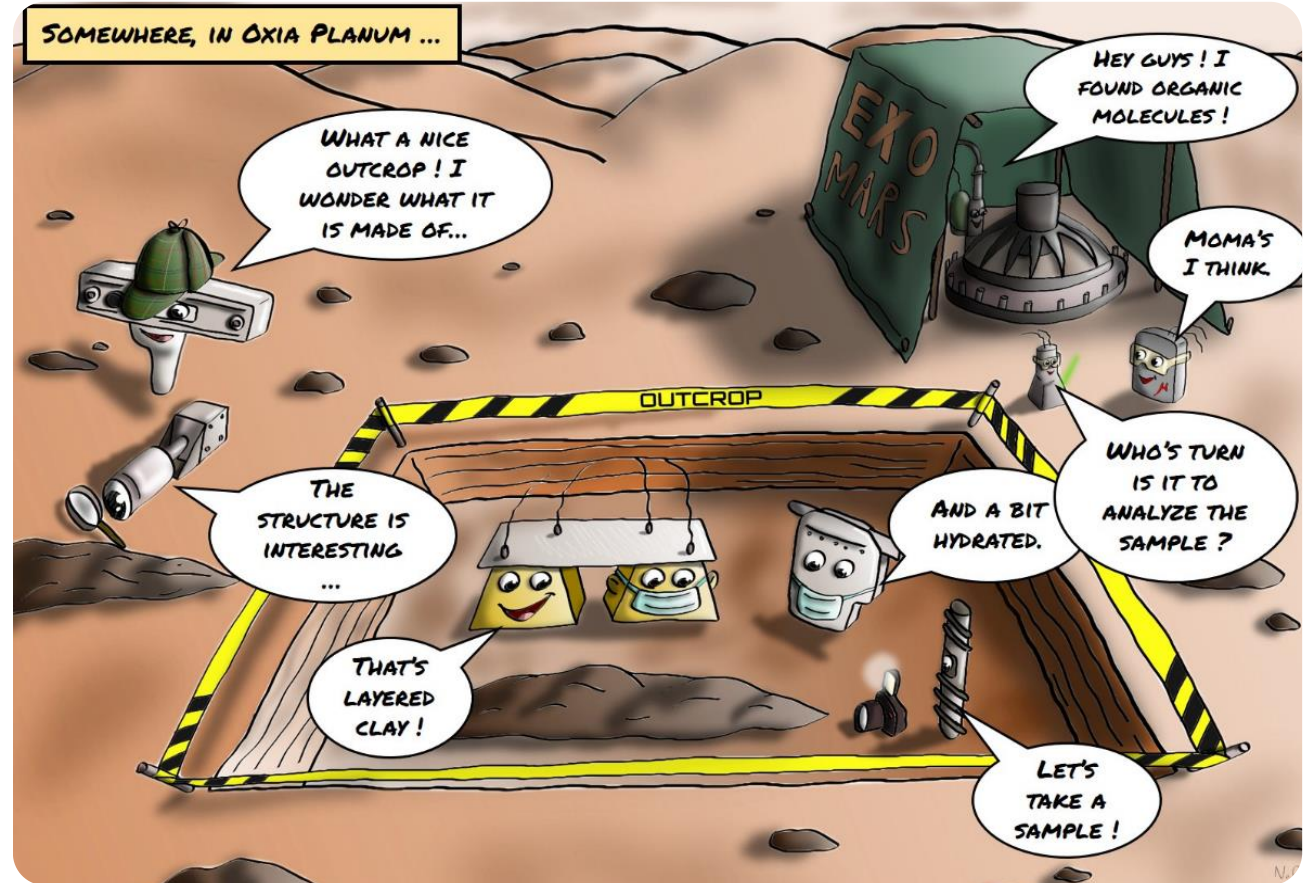
Rosalind Franklin Rover



Scientific objectives of the ExoMars mission :

- Study of the shallow subsurface (first meters)
- Seek for evidence of past life
- Study of the water distribution in the first few meters

Scientific instruments collaboration for the exploration of Oxia Planum subsurface

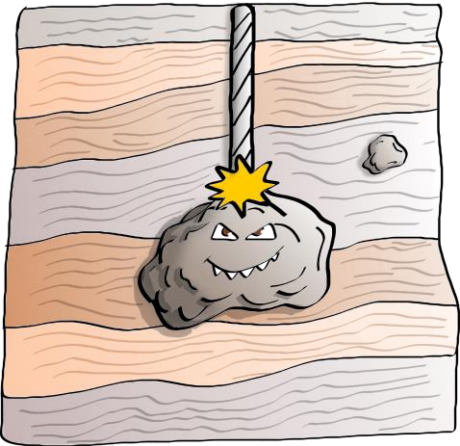


© Nicolas Oudart

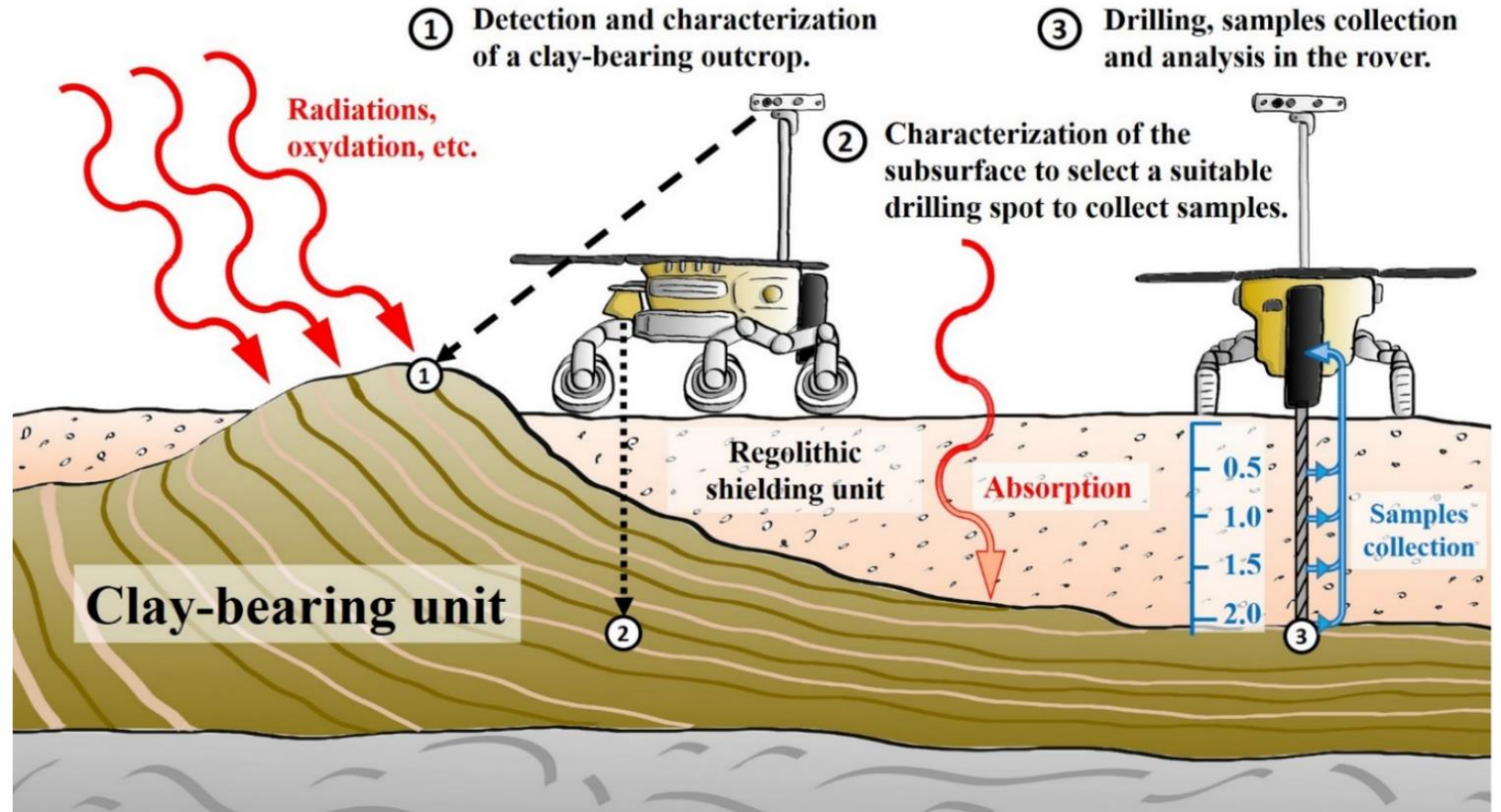


WISDOM OBJECTIVES IN EXOMARS MISSION

Identify interesting and **safe** sites for the drill



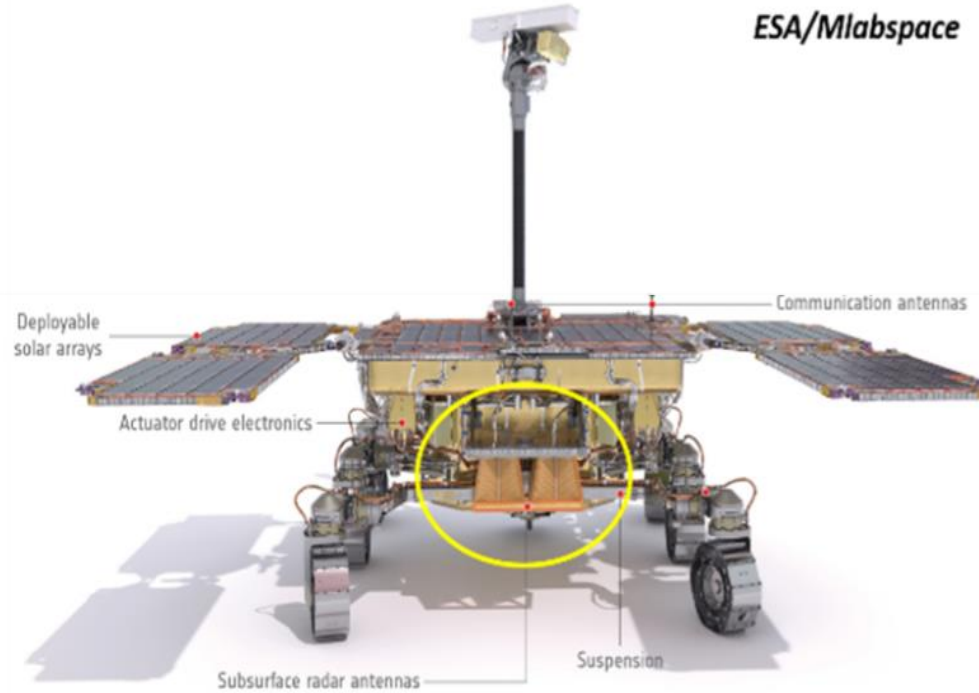
Understanding the geological history of the landing site :
Detect and characterize buried units (rocks, interfaces, ...)



© Nicolas Oudart



WISDOM INSTRUMENT



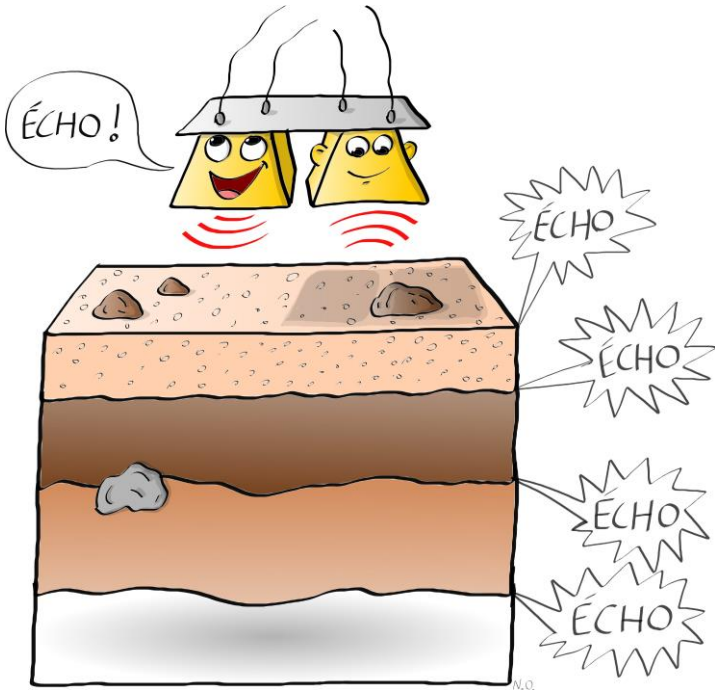
Miniaturized for space constraints in the Rover

15 cm

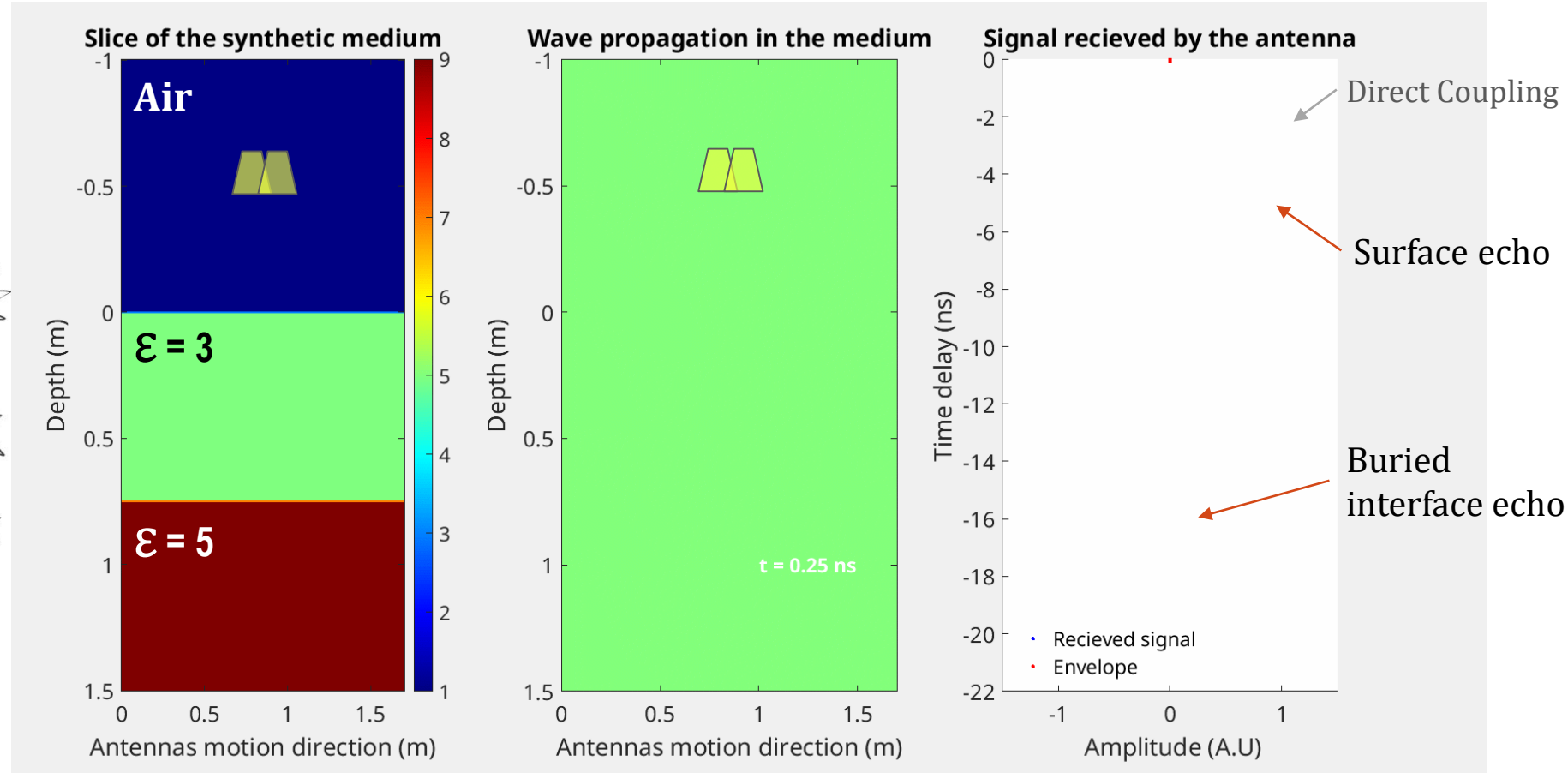
Radar design	Antennas design	Working Frequencies	Dynamic range	Power Consumption	Emitted Power	Mass (Antennas + Electronics)
Stepped Frequency Continuous wave	Vivaldi Antennas	0.5 - 3 GHz	~ 84 dB	12.5 W (peak)	1 mW	1.36 kg



GPR OPERATIONS AND DATA PRODUCTS



© Nicolas Oudart

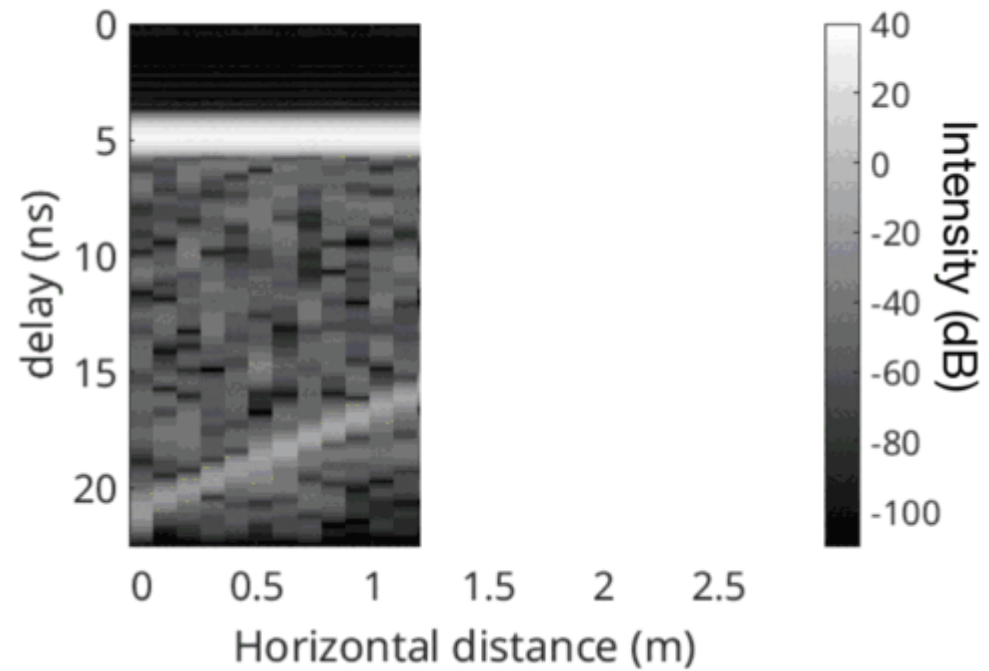
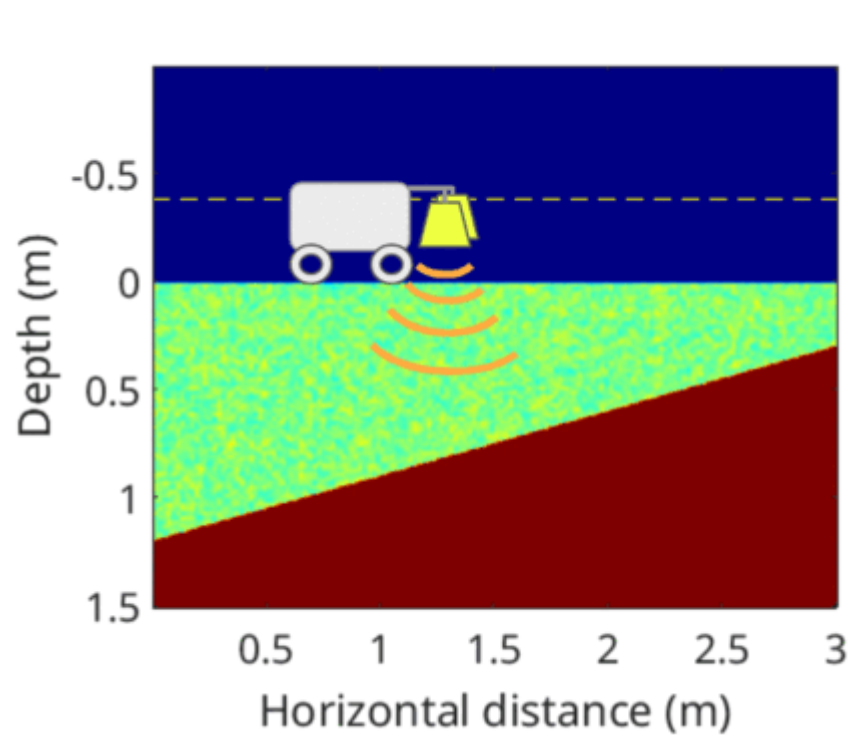




WISDOM
GPR



GPR OPERATIONS AND RADARGRAM



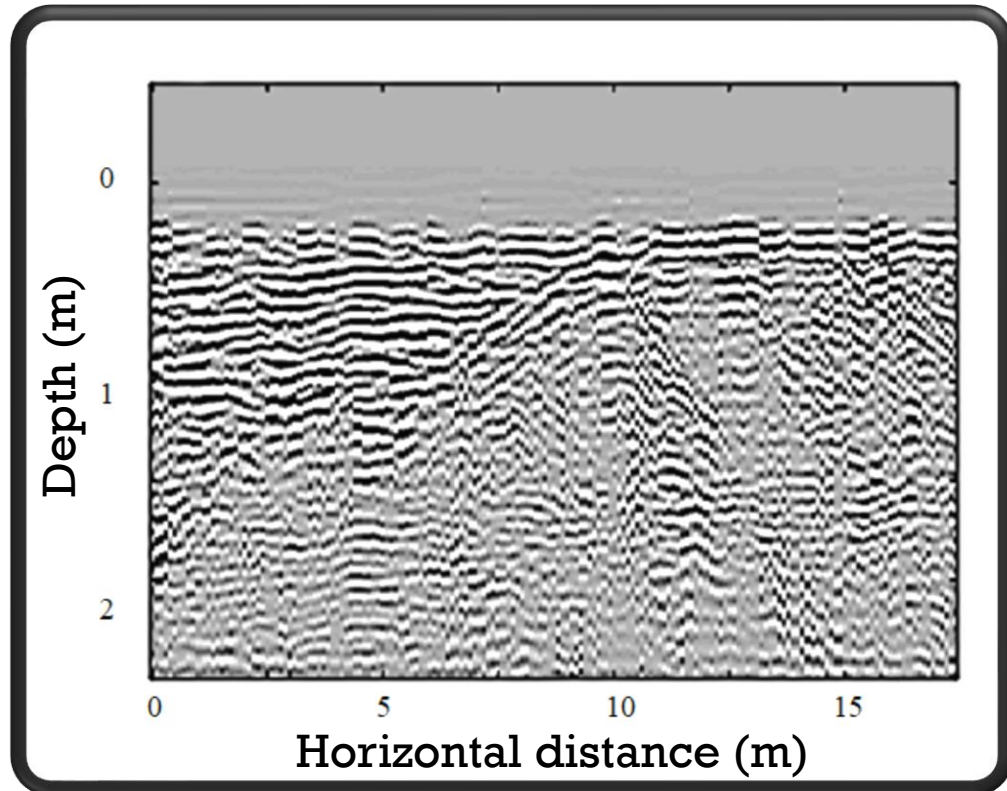


WISDOM
GPR

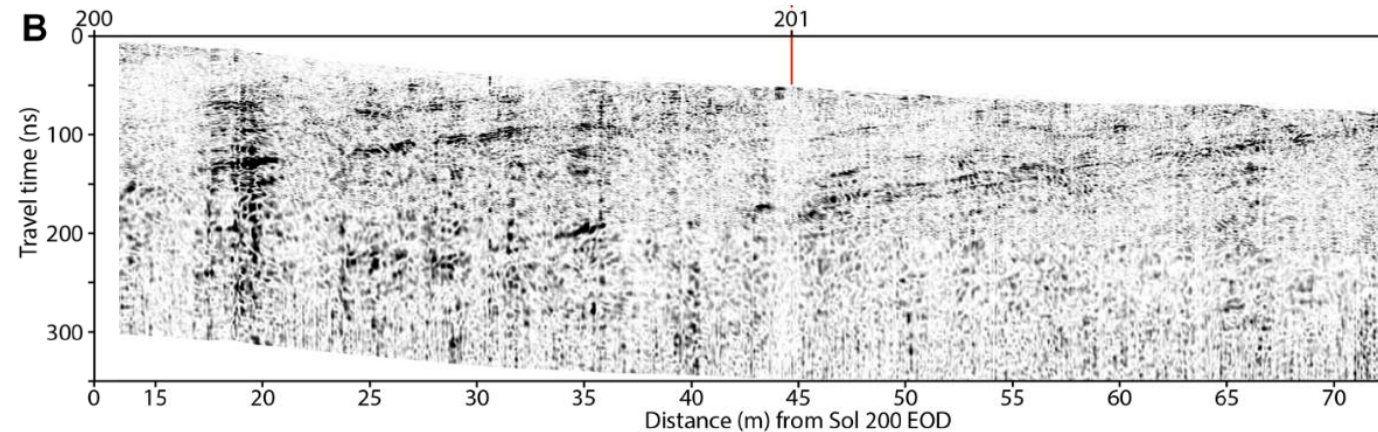


GPR OPERATIONS AND DATA PRODUCTS

Earth, Field test campaign in Colorado
Provençal (France) – WISDOM Prototype



Mars, Jezero Crater – RIMFAX GPR onboard
Perseverance Rover

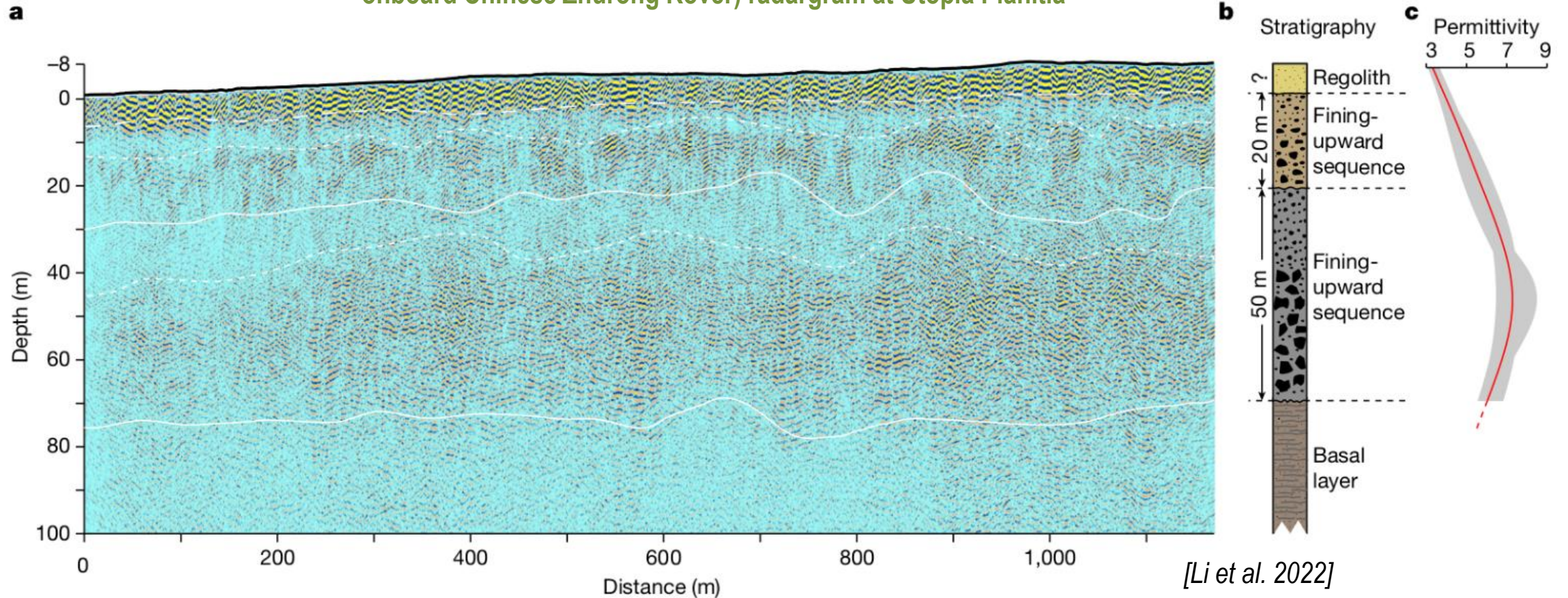


[Hamran S. 2022]



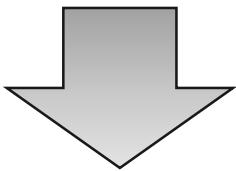
OBJECTIVE : CHARACTERIZING HETEROGENEOUS MEDIA

Radargrams acquired in natural environments on Mars in 2022 from RoPeR (GPR onboard Chinese Zhurong Rover) radargram at Utopia Planitia

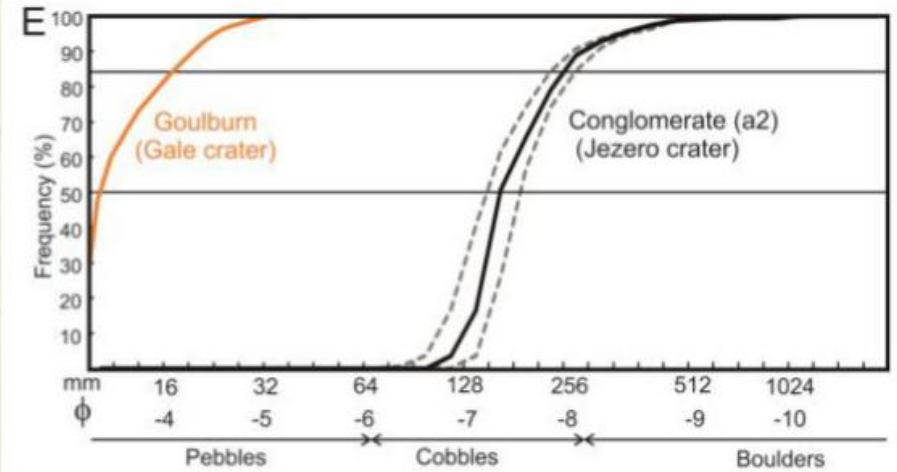


MARTIAN PICTURE FROM ANOTHER ROVER ON MARS

Scatterer size distribution



Clues about past
hydraulic / eolian
activities



[Mangold et al. 2021]

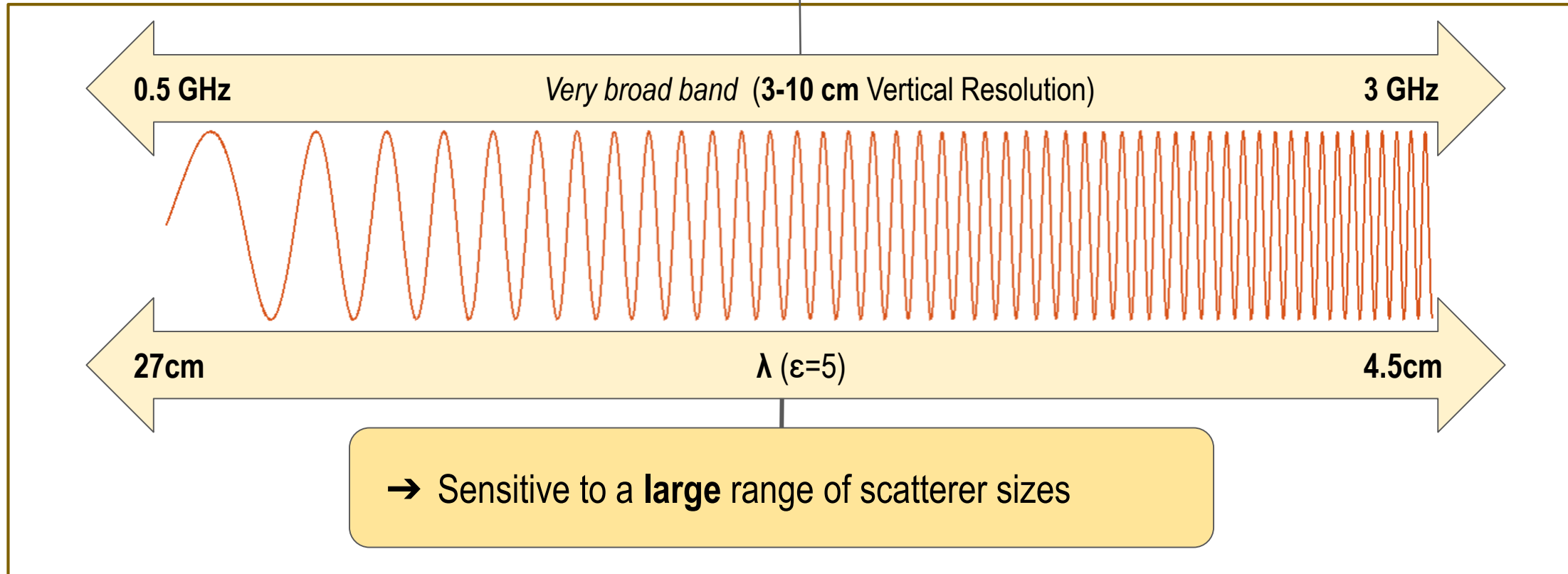


WISDOM PERFORMANCES

WISDOM is a *polarimetric step frequency* Ground Penetrating Radar

Central frequency = 1.75 GHz

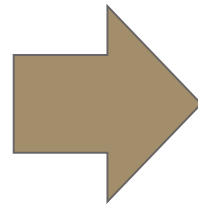
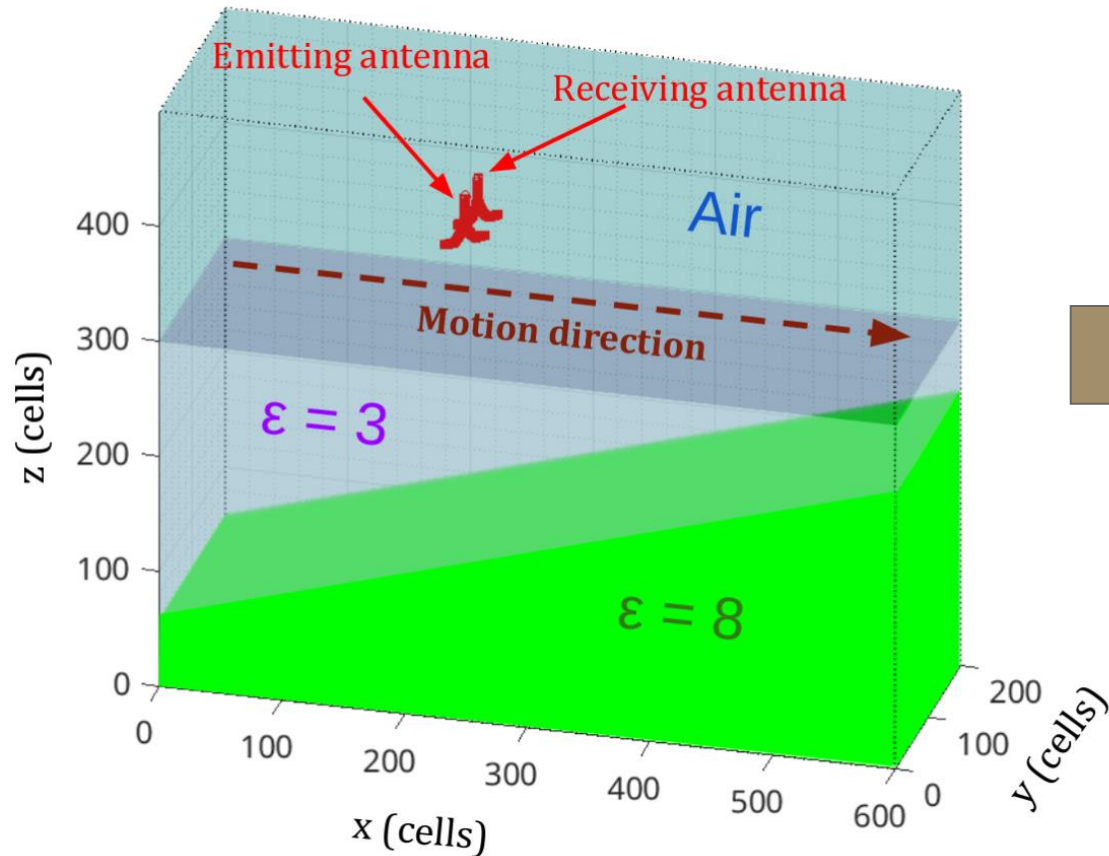
→ Penetration depth of ***few meters***



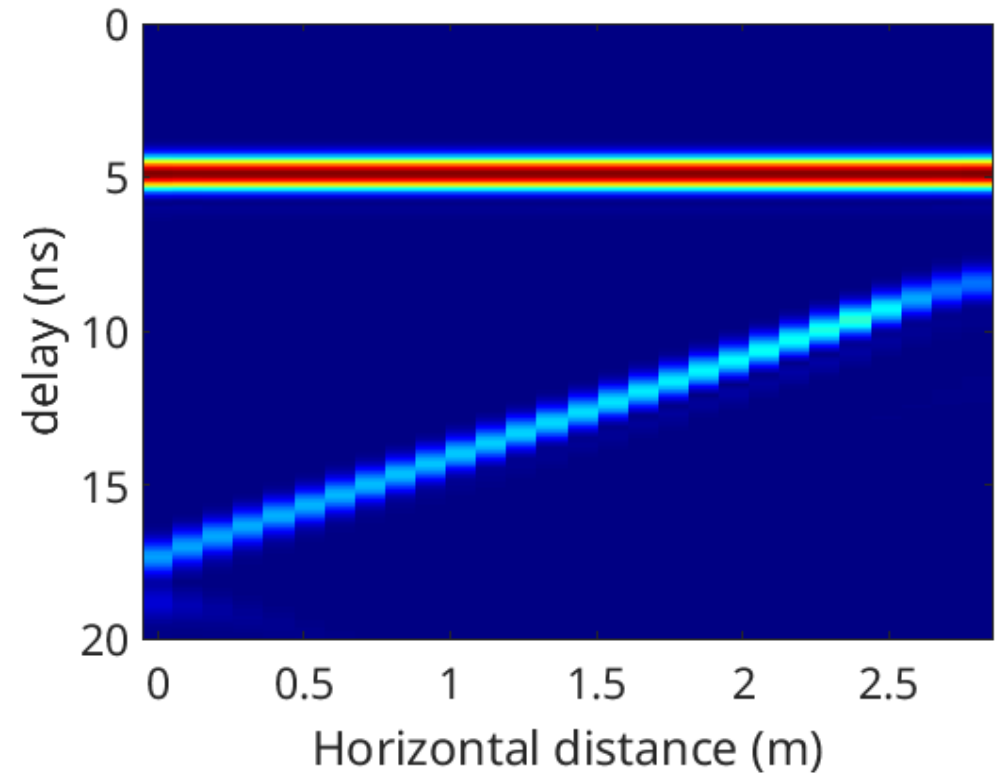
3D SIMULATIONS WITH FDTD CODE

Numerical simulations are performed with *Finite Domain Time Difference* method (TEMSEI-FD, XLIM, France)

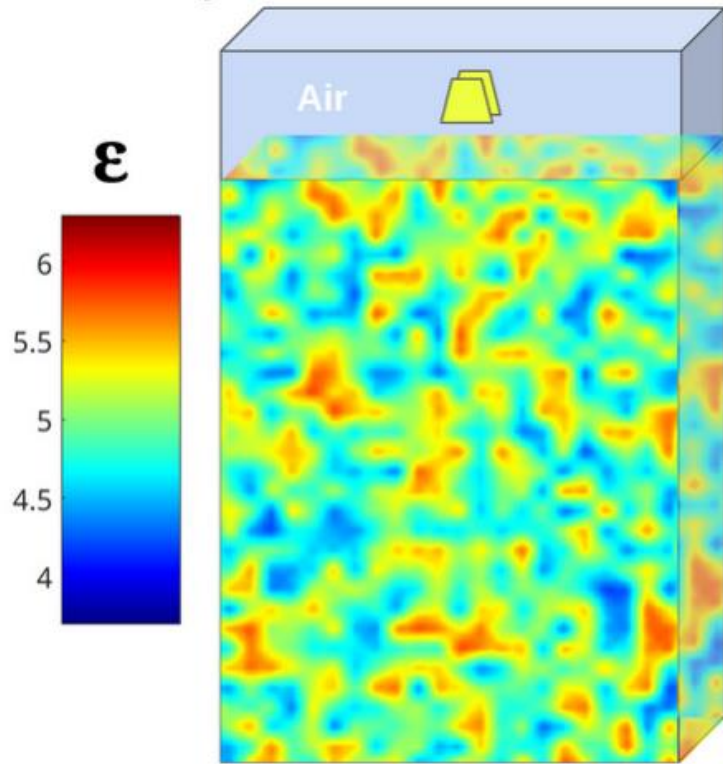
Computing synthetic volume



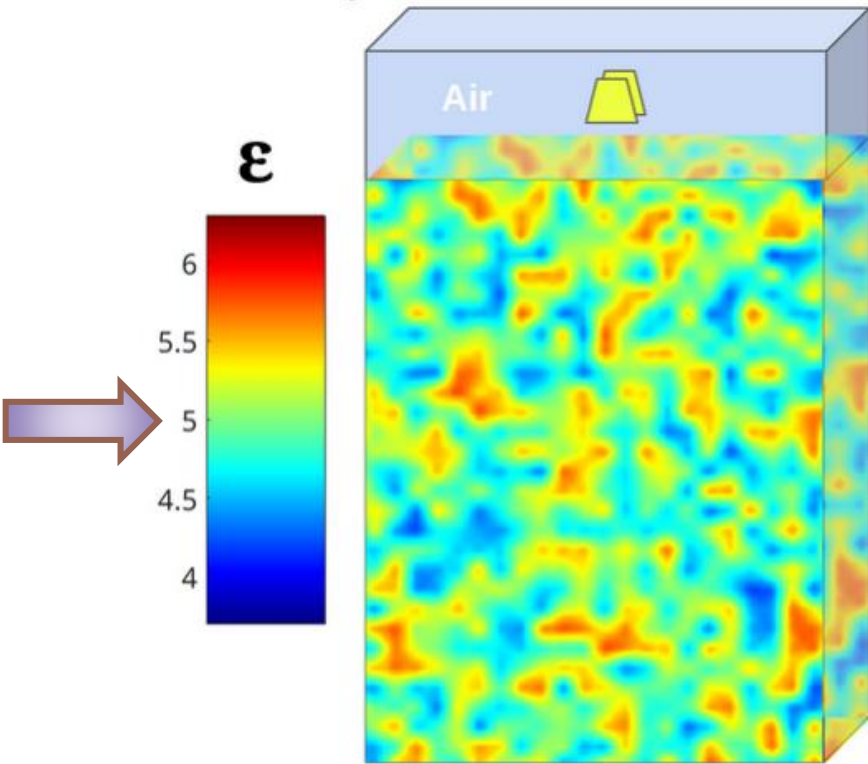
WISDOM synthetic radargram



3D SIMULATIONS OVER HETEROGENEOUS MEDIA



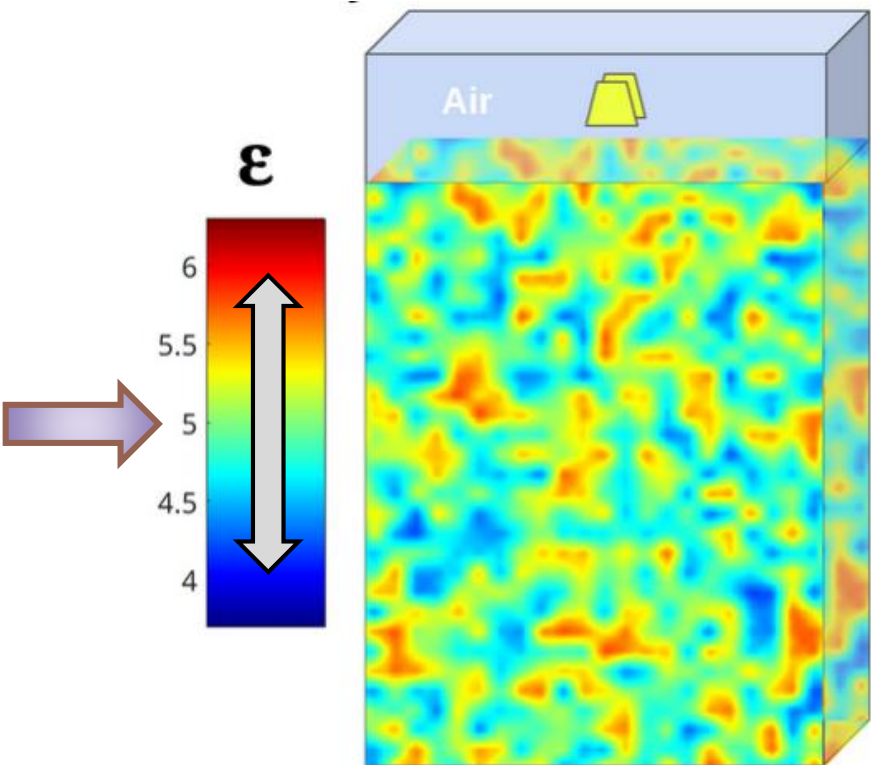
3D SIMULATIONS OVER HETEROGENEOUS MEDIA



→ The mean permittivity $\bar{\epsilon}$



3D SIMULATIONS OVER HETEROGENEOUS MEDIA

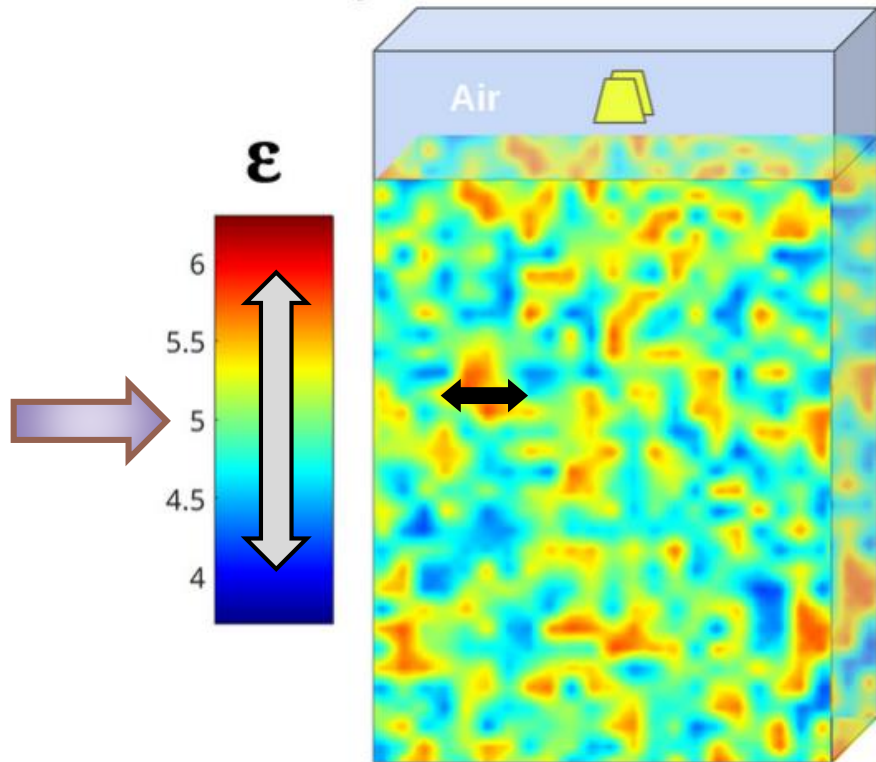


→ The mean permittivity $\bar{\epsilon}$

→ The standard deviation of permittivity $\Delta\epsilon$



3D SIMULATIONS OVER HETEROGENEOUS MEDIA



→ The mean permittivity $\bar{\epsilon}$

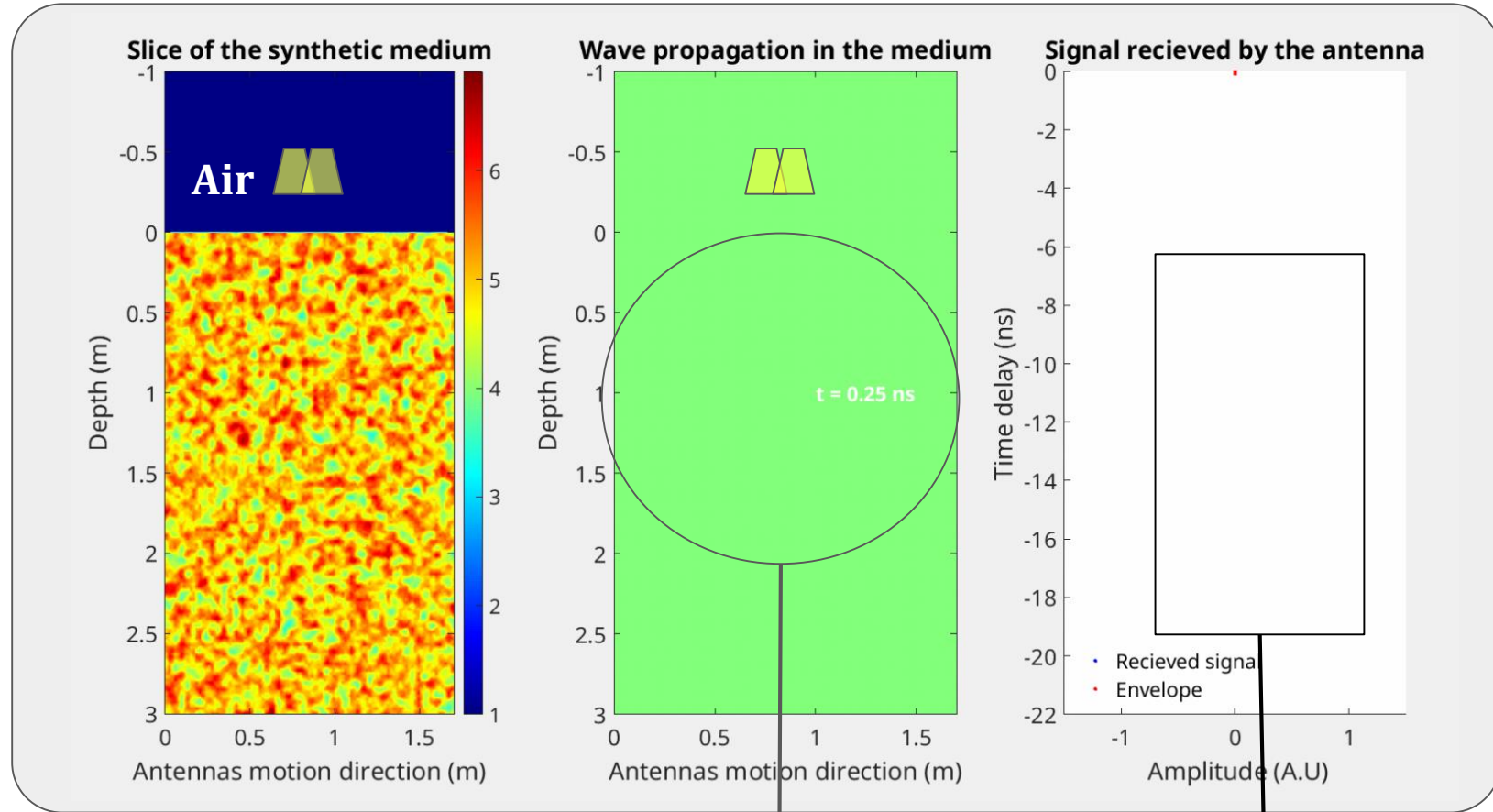
→ The standard deviation of permittivity $\Delta\epsilon$

→ The typical size of heterogeneities L



3D SIMULATIONS OVER HETEROGENEOUS MEDIA

$\epsilon = 5$
 $\Delta\epsilon = 0.3$
 $L = 3.2 \text{ cm}$



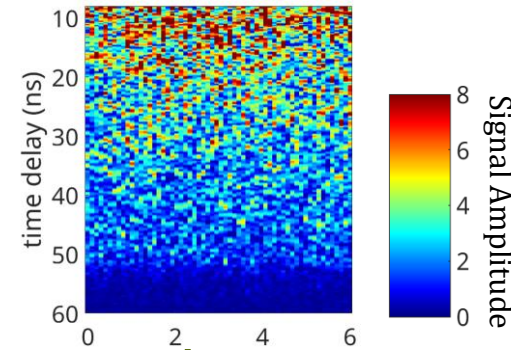
Volume scattering

Volume backscatter



DIVISION IN SUB-FREQUENCY BANDS

Full WISDOM frequency band

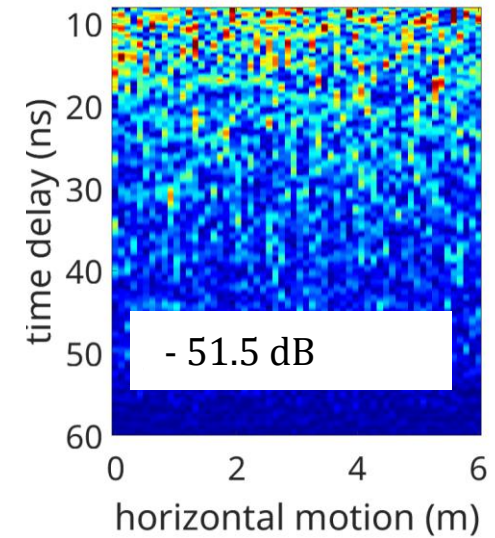
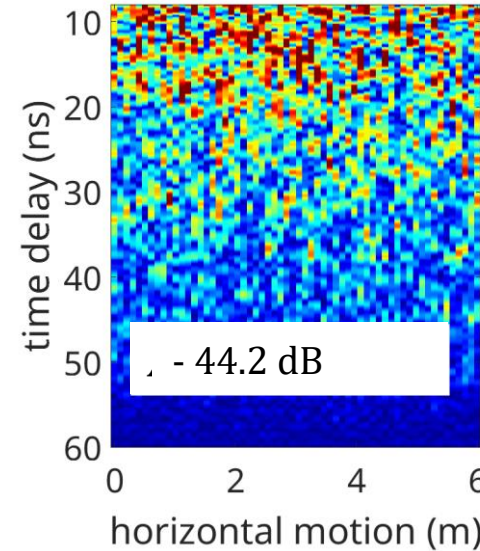
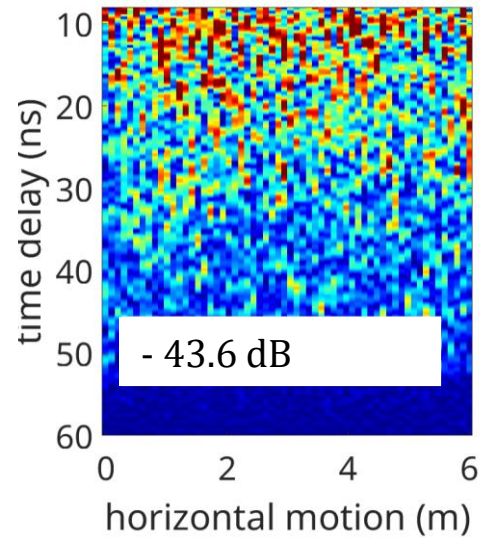
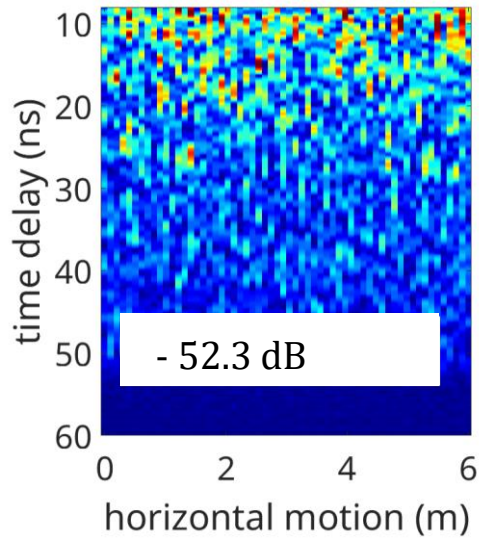


$\lambda \approx 17$ cm
0.5 - 1.12 GHz

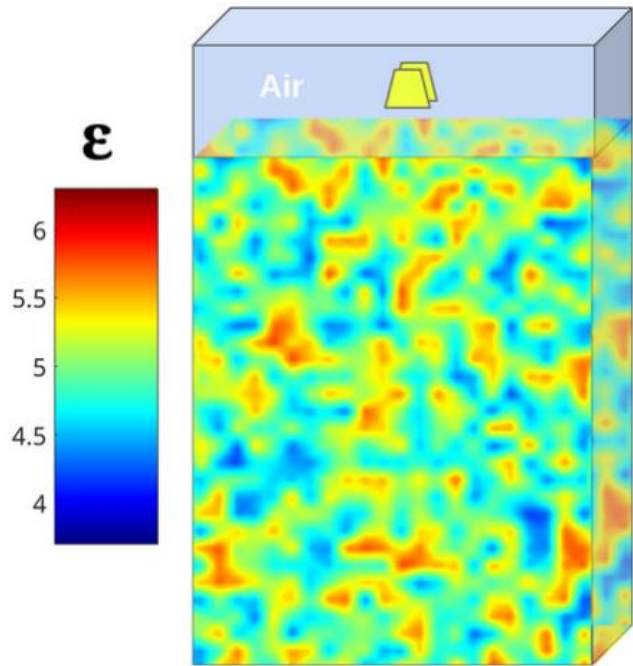
$\lambda \approx 9.3$ cm
1.12 - 1.75 GHz

$\lambda \approx 6.5$ cm
1.75 - 2.37 GHz

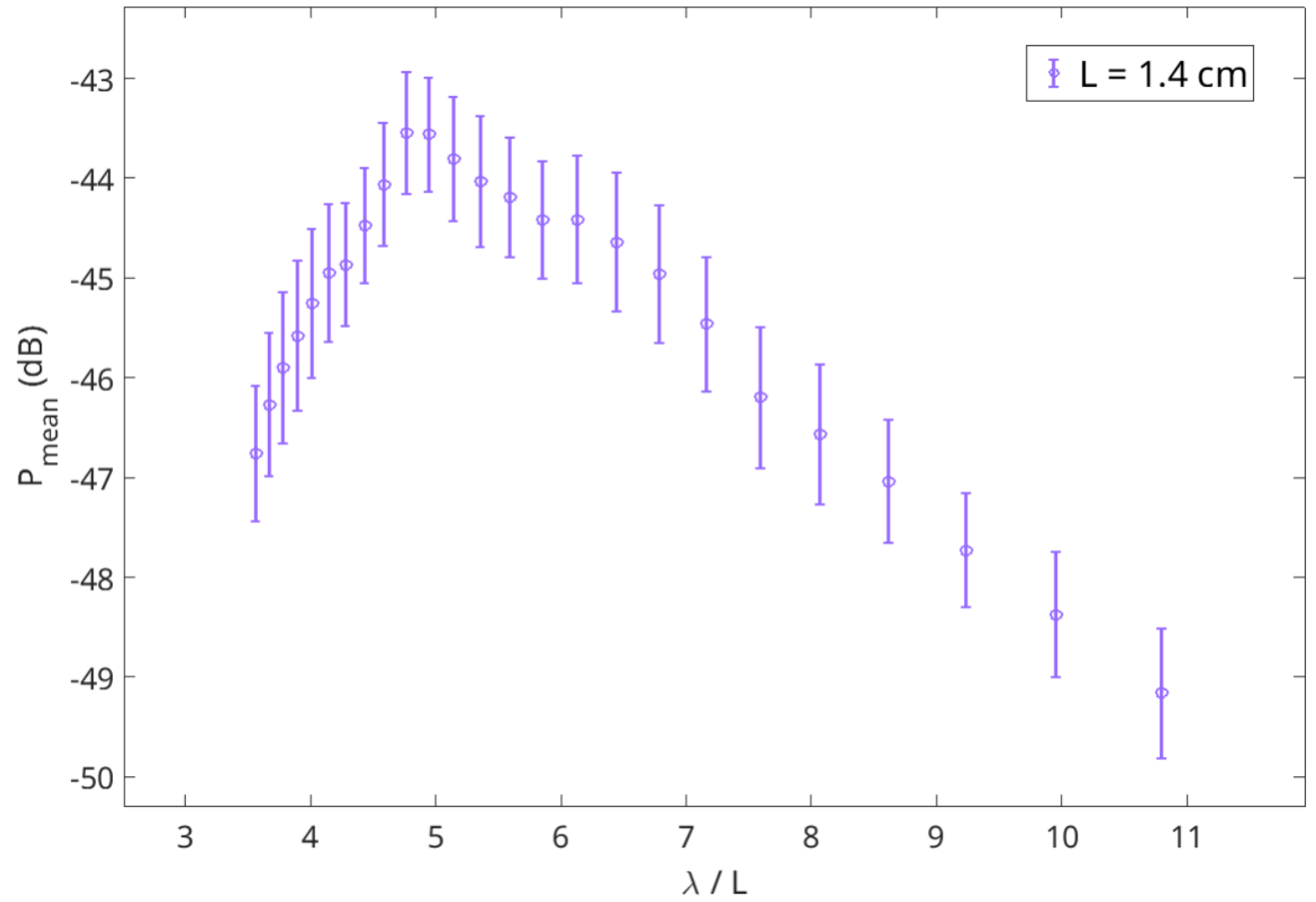
$\lambda \approx 4.9$ cm
2.37 - 3 GHz



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES



$$\begin{aligned}\epsilon &= 5 \\ \Delta\epsilon &= 0.3 \\ L &= 1.4 \text{ cm}\end{aligned}$$

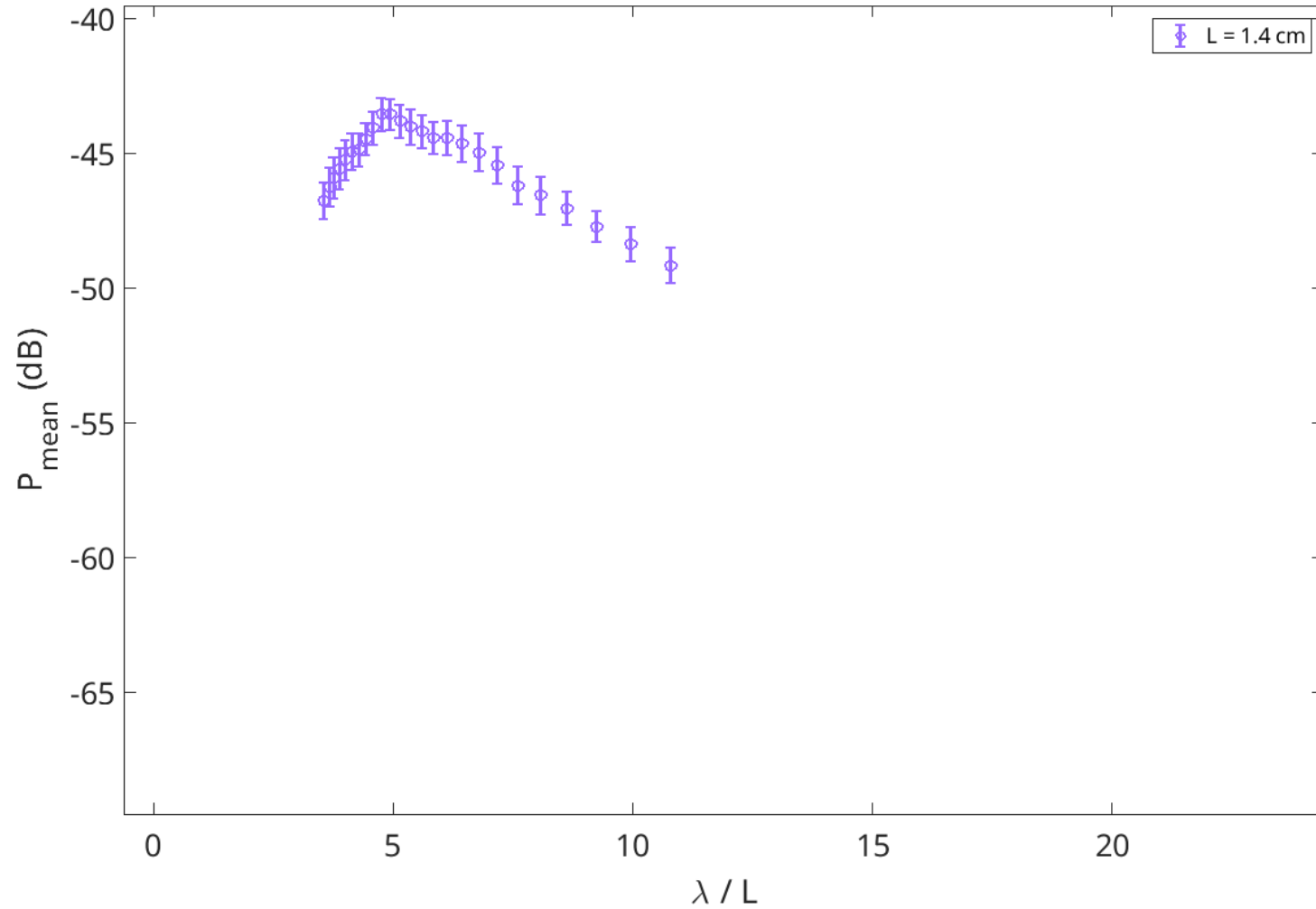




RESULTS



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES

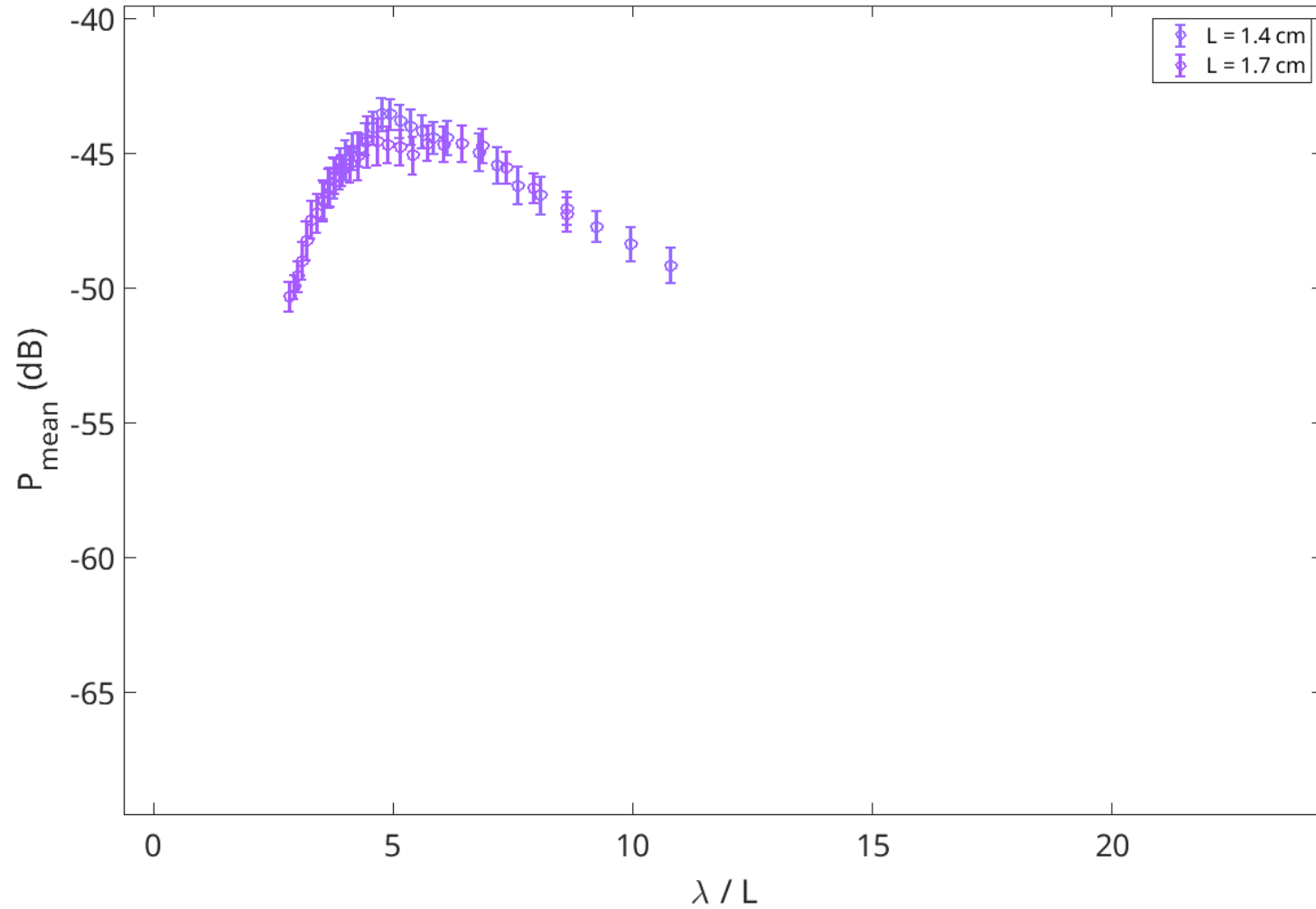




RESULTS



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES

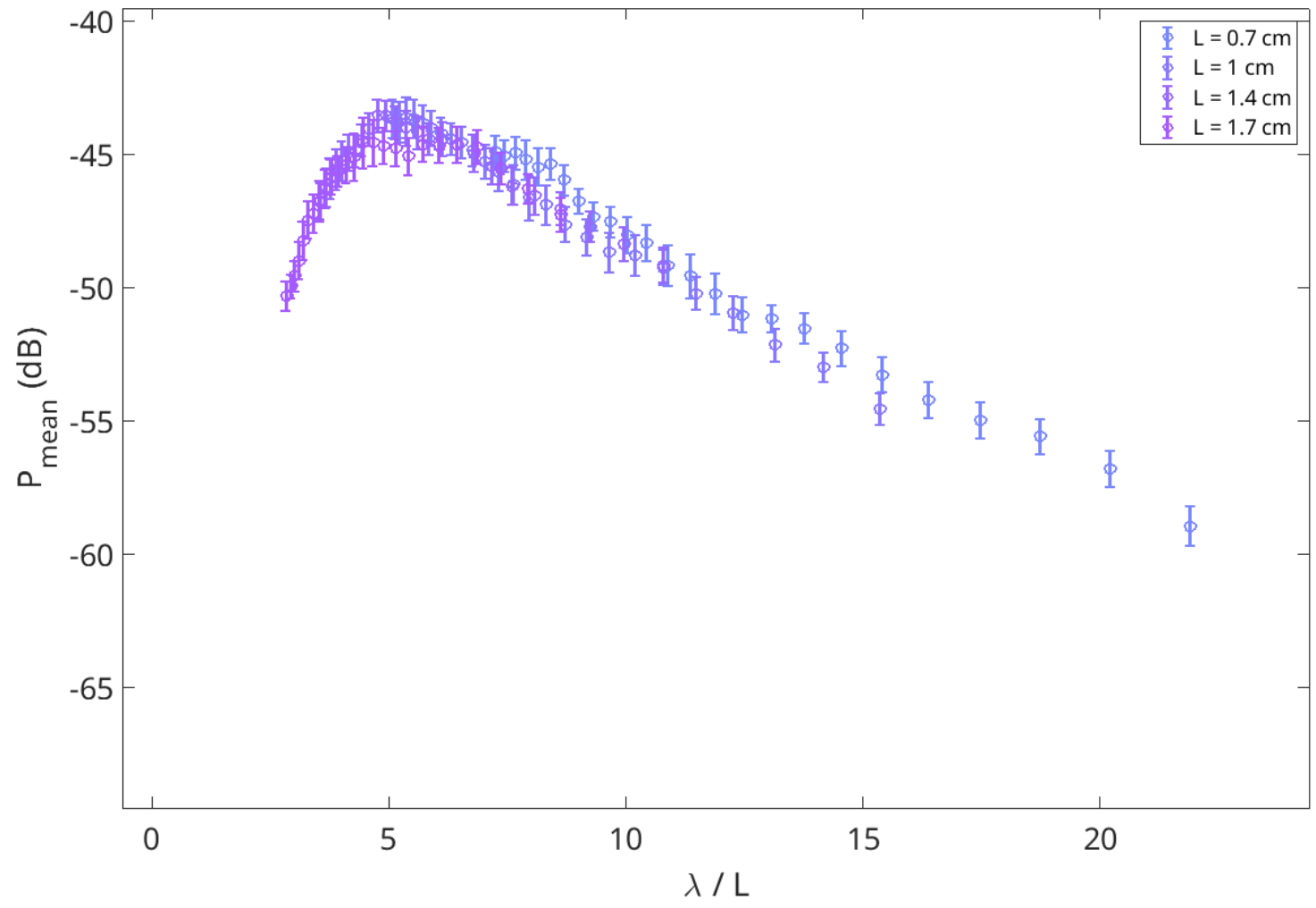




RESULTS



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES

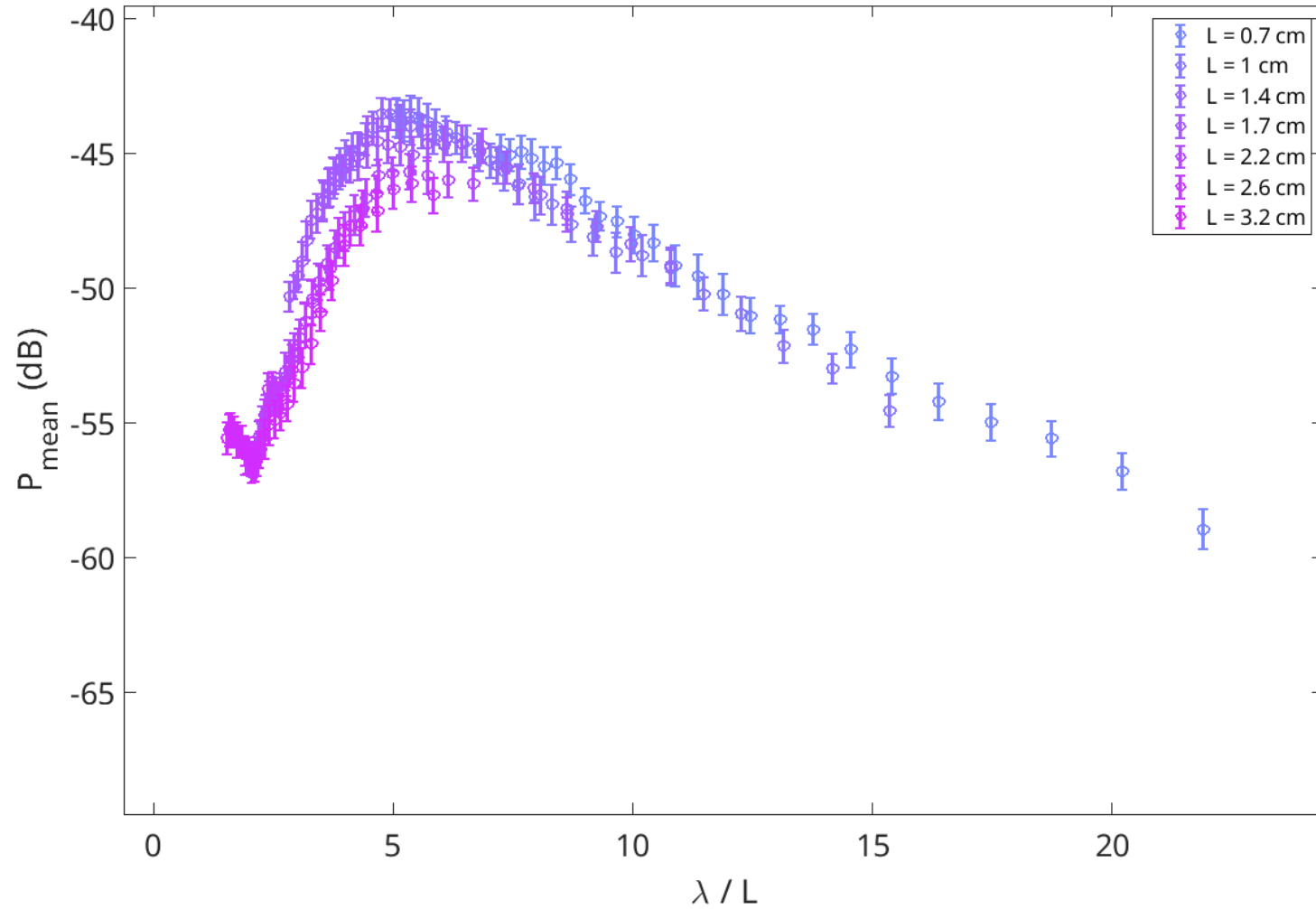




RESULTS



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES

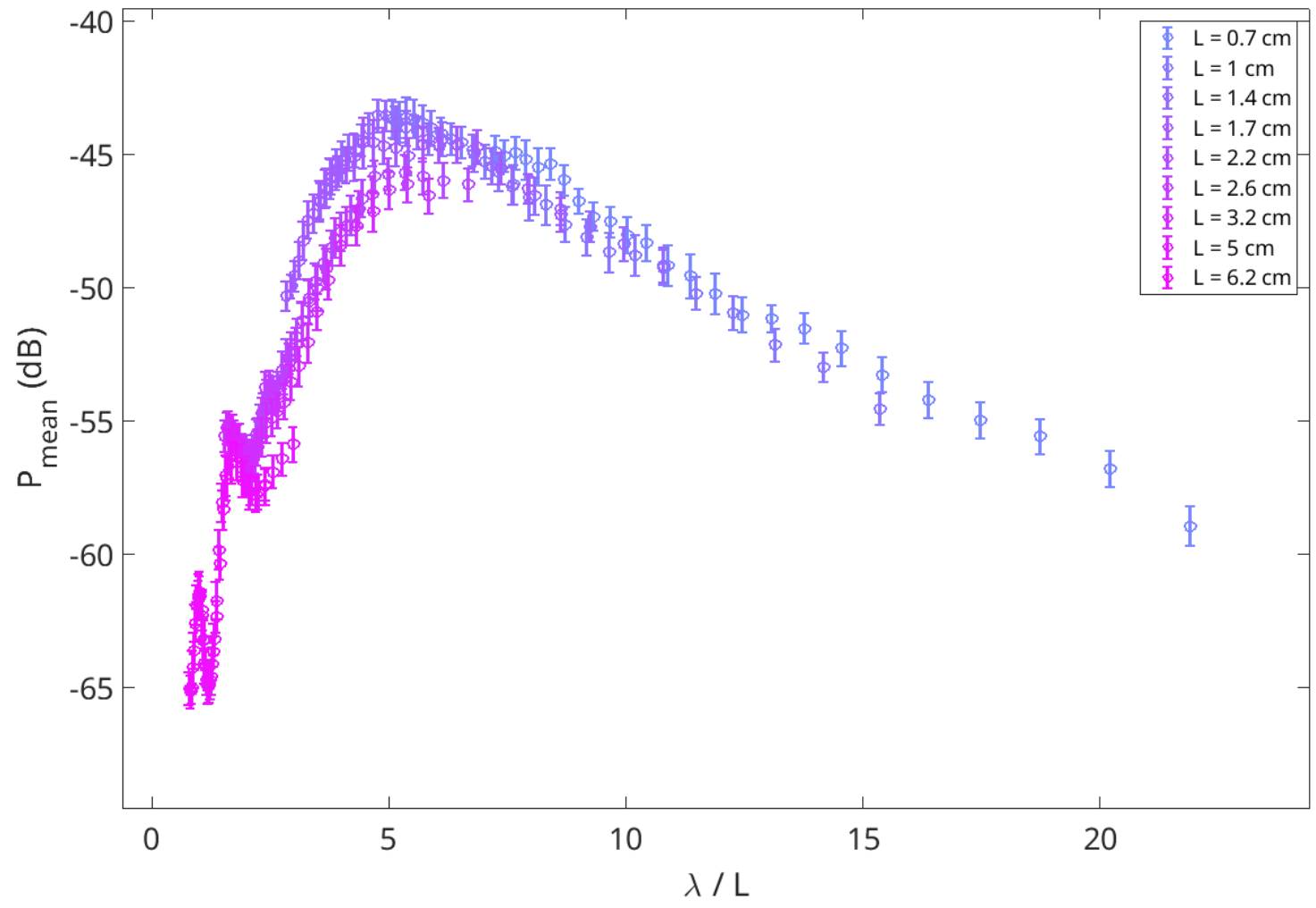




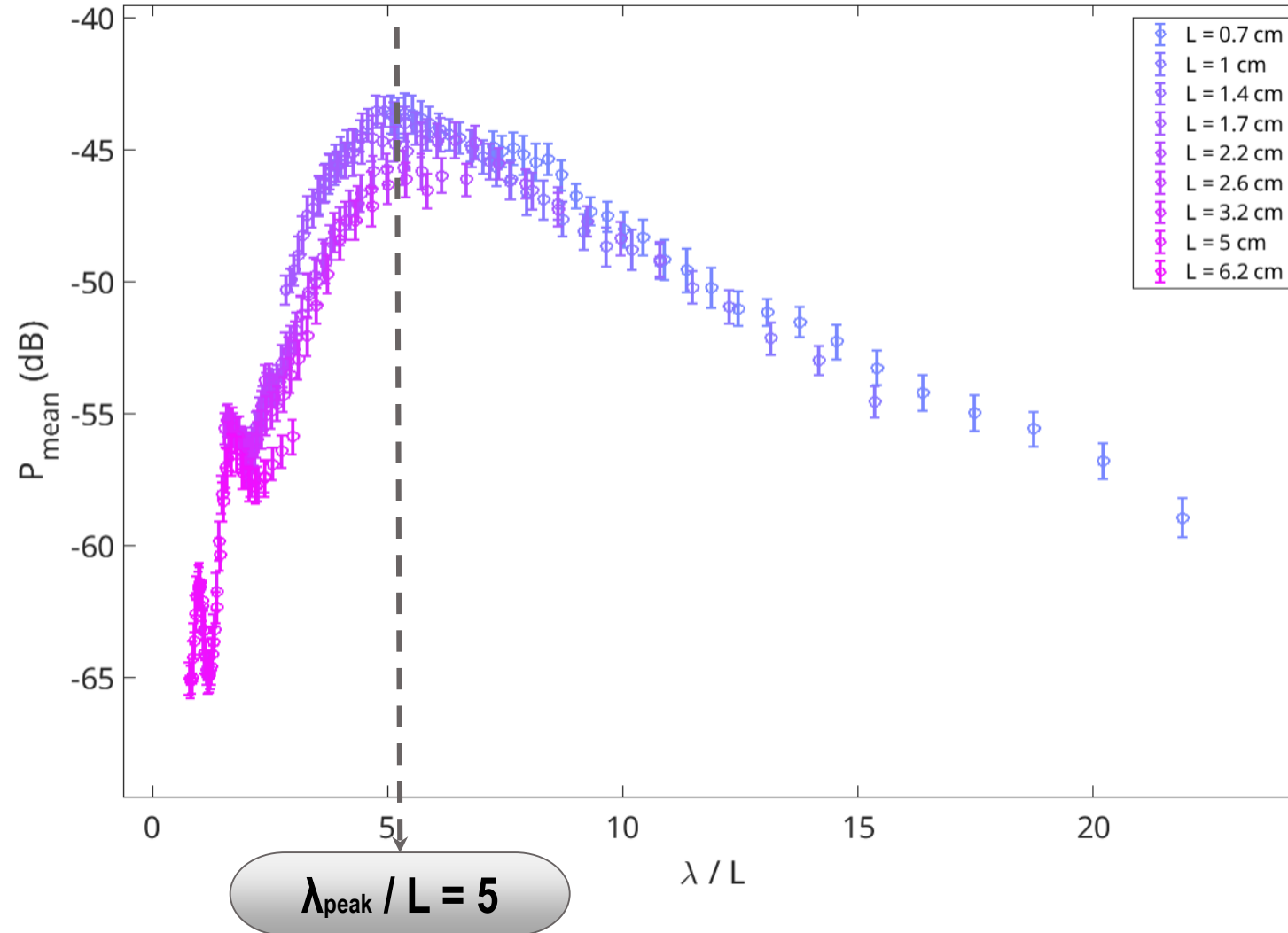
RESULTS



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES



RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES



The maximum is reached for :

$$\lambda_{\text{peak}} / L = 5$$

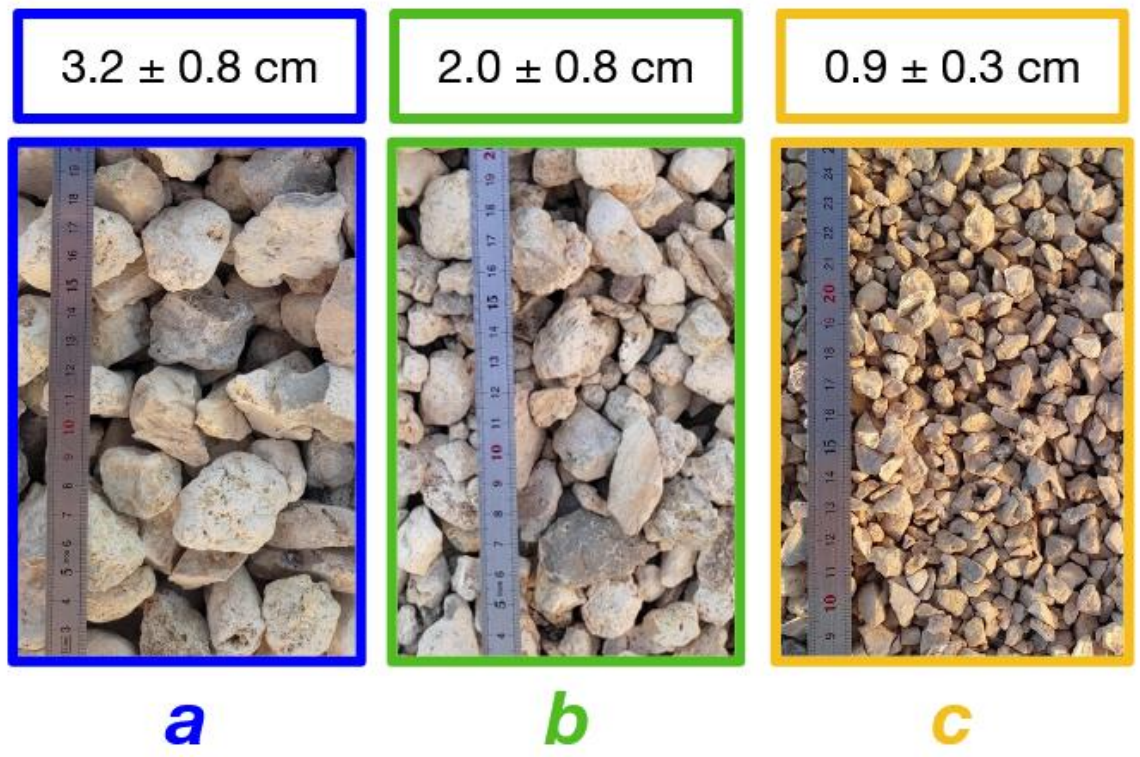
Retrieval of L ? Determine

λ_{peak} and then :

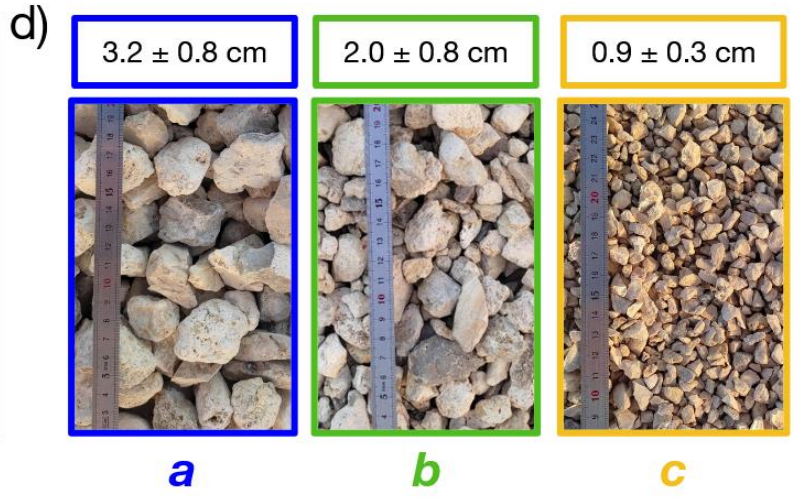
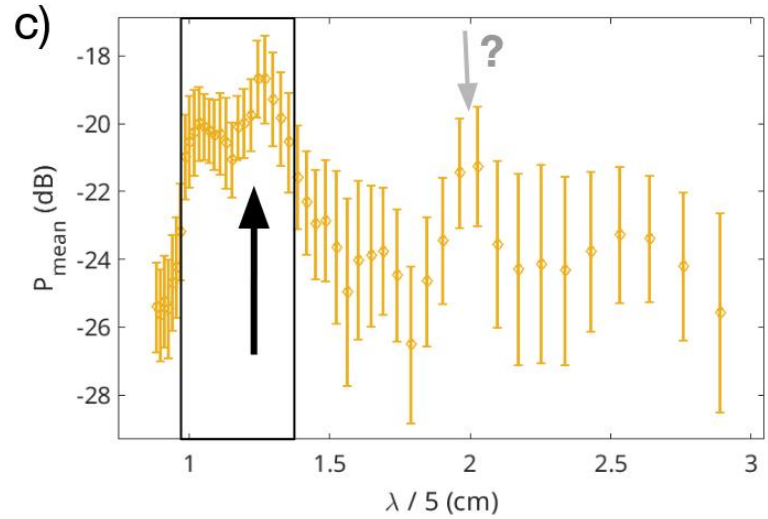
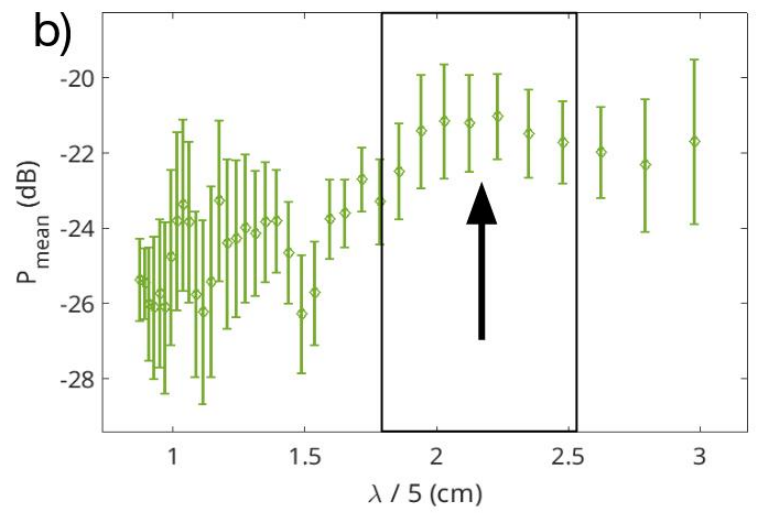
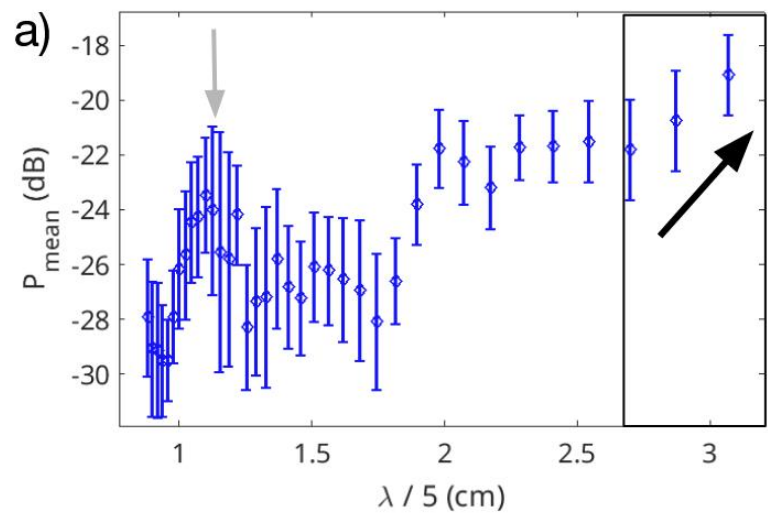
$$L = \lambda_{\text{peak}} / 5$$



VALIDATION ON EXPERIMENTAL DATA



VALIDATION ON EXPERIMENTAL DATA



CONCLUSION & PERSPECTIVES

- ❖ We successfully elaborated a method to retrieve the **typical size of heterogeneities** which provides information on the geological context of the site.
- ❖ The method presented still applies for different permittivity distributions and in most lossy media
- ❖ **This approach can be applied to data from any broad-band GPR**
- ❖ Preliminary analysis show promising results not only in co-polarization but also in **cross-polarization** configuration.
- ❖ The method has been preliminarily applied on real data, acquired on controlled environment.
- ❖ Ongoing study shows similar results if the scatterers are spheres.

Perspectives :

- The method could be tested on multi-layer heterogeneous subsurfaces
- Further validations on **controlled and documented environments** are warranted
- Can be applied on data collected on **Martian** and **lunar** analogs





Thank you !

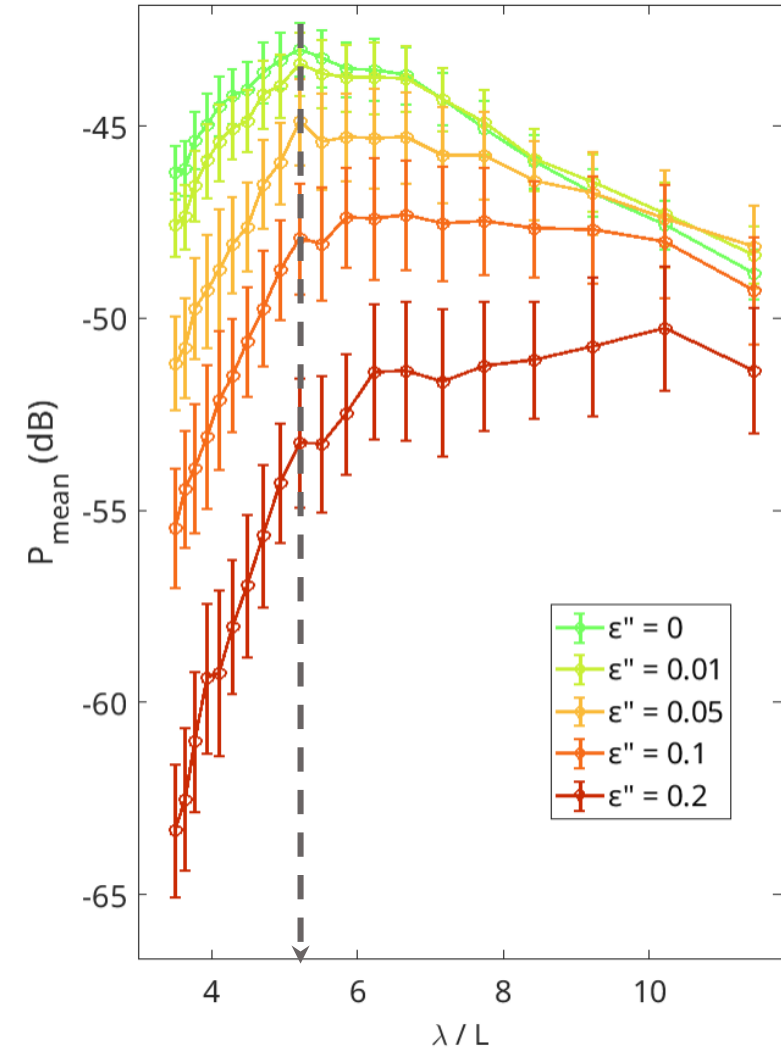
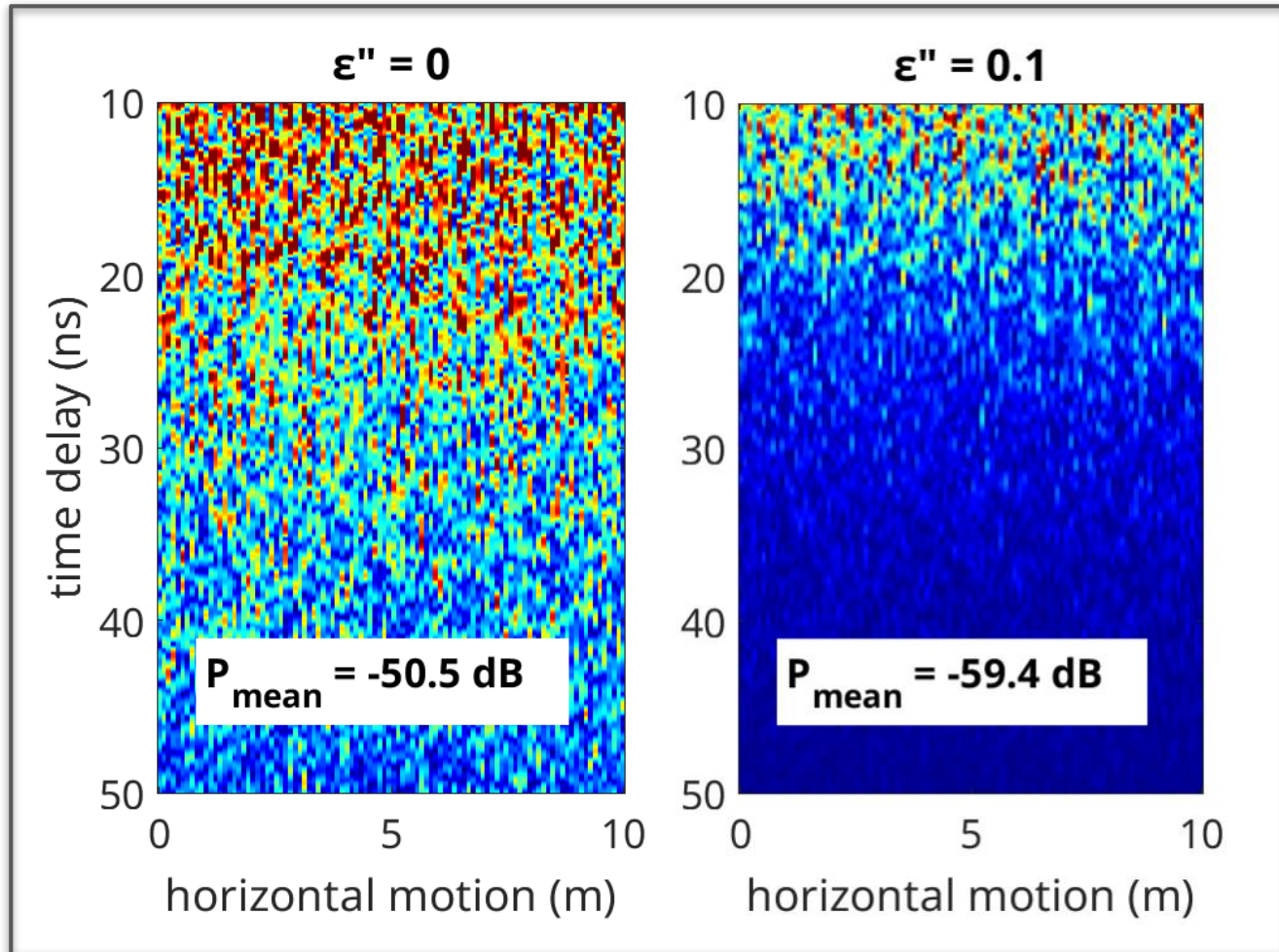
Contact : emile.brighi@latmos.ipsl.fr



IMPACT OF DIELECTRIC LOSSES

$$(\epsilon = \epsilon' - j\epsilon'')$$

$L = 1.4 \text{ cm}$



IMPACT OF DIELECTRIC LOSSES

Martian and Lunar soil analogs (color indicates the typical ϵ'' values of the material) :

Basalt

Martian Dust

Dry sandstones

Moon regolith

Moon regolith

Mars JSC-1

Mars Jezero Crater

Dry clays

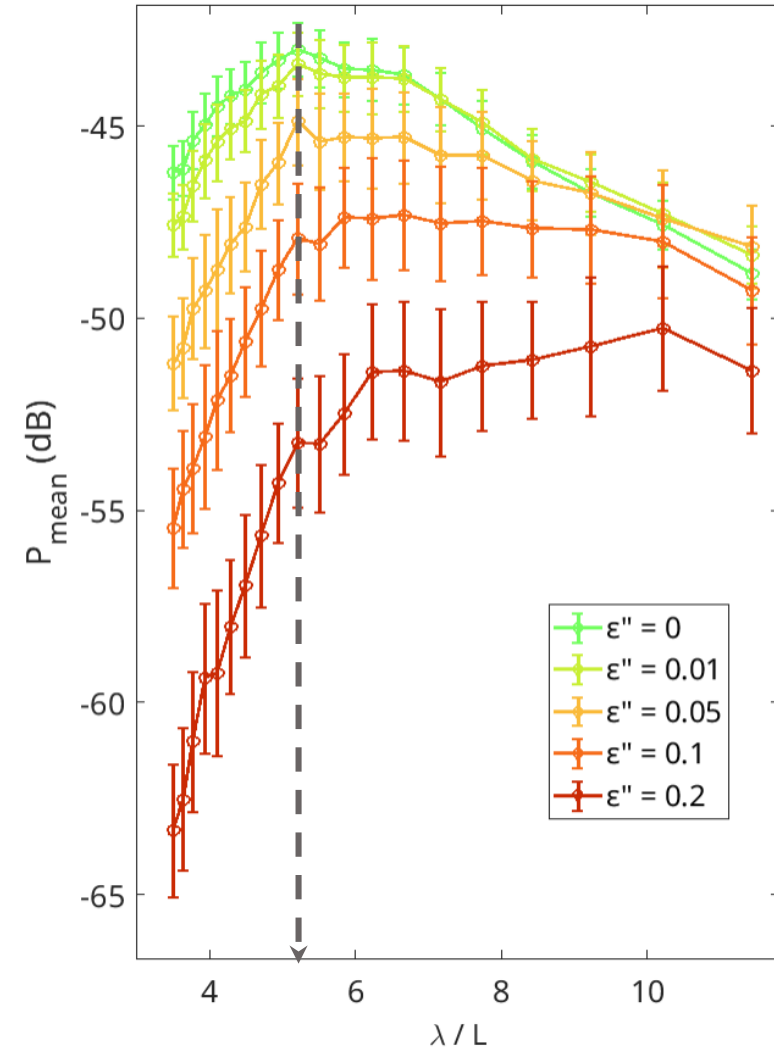
Carbon Red Clay



RESULTS



$L = 1.4$ cm



IMPACT OF DIELECTRIC LOSSES

Martian and Lunar soil analogs (color indicates the typical ϵ'' values of the material) :

Basalt

Martian Dust

Dry sandstones

→ Moon regolith (Returned samples)

→ Moon regolith (GPR estimates)

Mars JSC-1

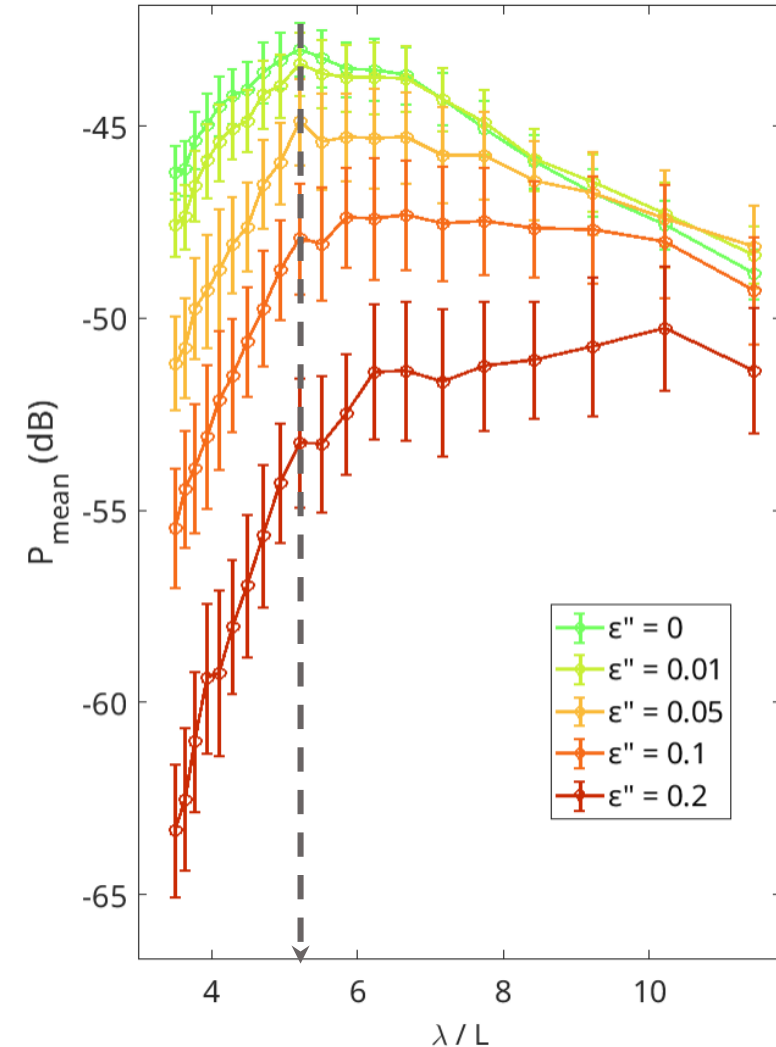
→ Mars Jezero Crater (RIMFAX estimates)

Dry clays

Carbon Red Clay



L = 1.4 cm



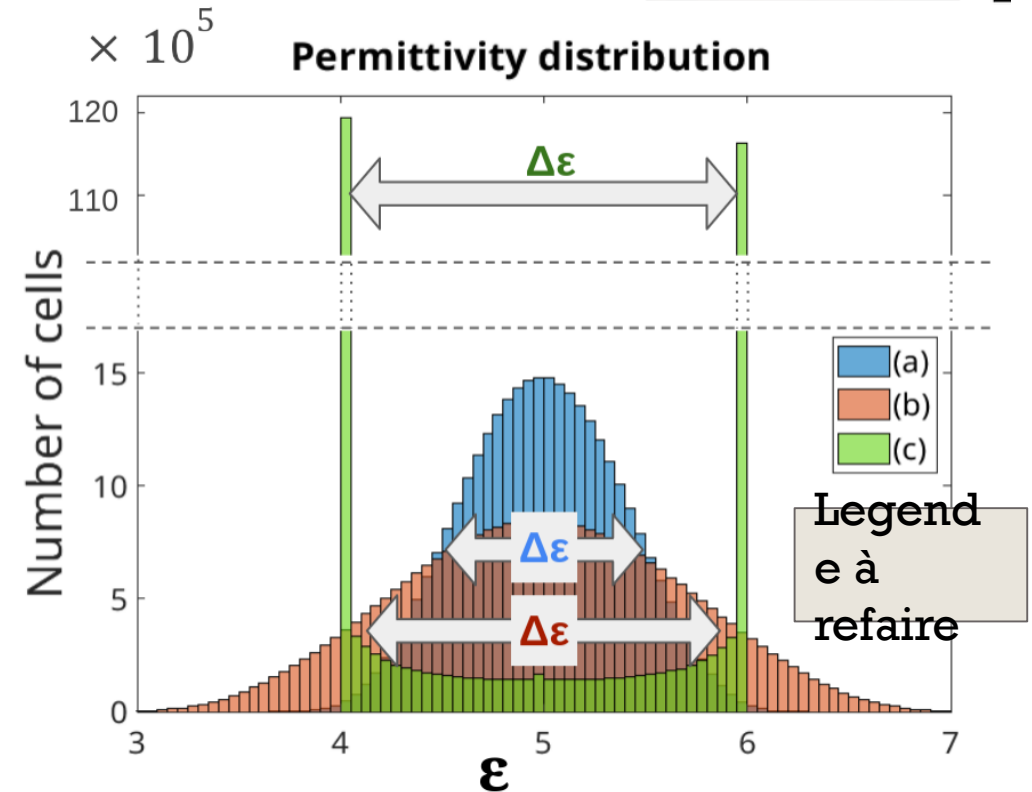
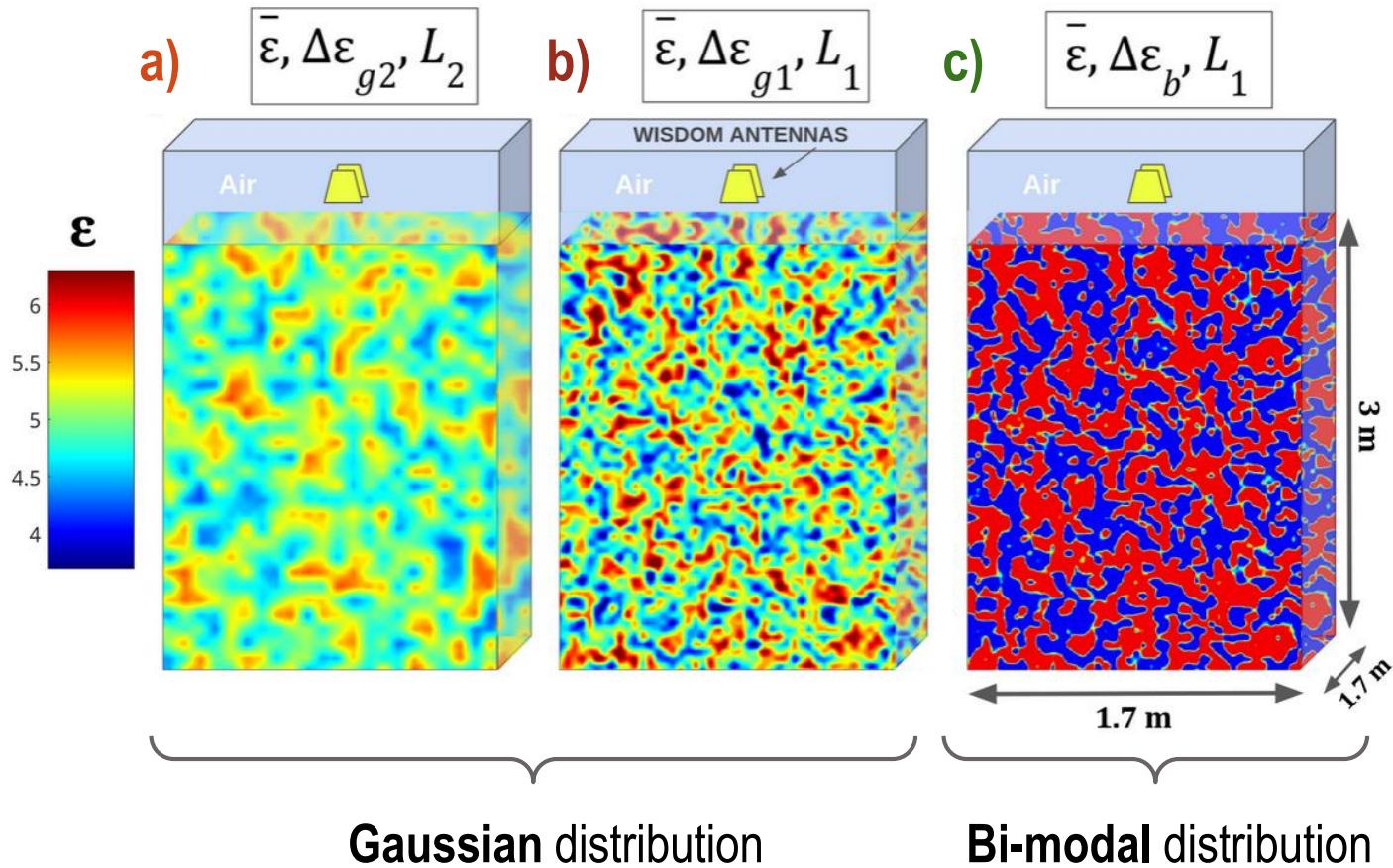
ANNEXE

Parameter of the instrument	Typical values for WISDOM
Frequency Bandwidth	0.5 to 3 GHz (2.5 GHz)
Frequency steps	2.5 MHz (1001 steps)
Repetition time	200 μ s
Transmitted power	0 dBm
Noise factor	8 dB
Receiving channel gain (adjustable)	-7 to 25 dB
Antenna gain	1 to 8 dB
Effective range of the ADC (16 bits)	\sim 84 dB
IFT gain	27 dB
Gain for 10 coherent additions	10 dB (see (Hervé, 2018))
Transmitter and receiver efficiency	0.1

[V. Ciarletti (2017) ; N. Oudart (2021)]



3D SIMULATIONS OVER HETEROGENEOUS MEDIA

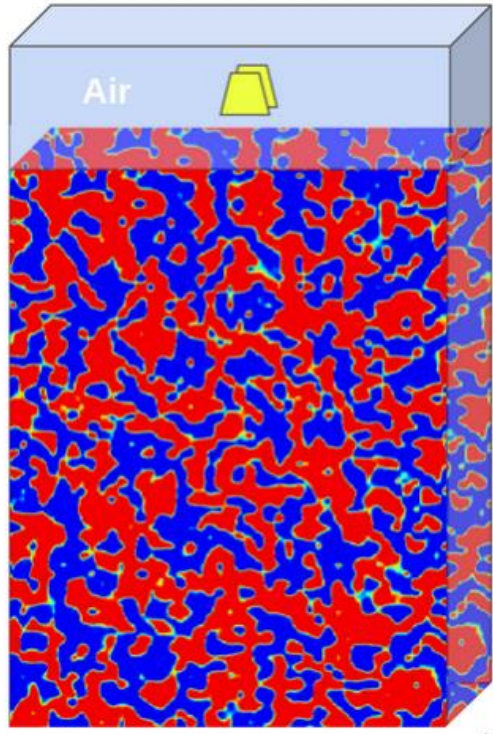


RETRIEVAL OF THE TYPICAL SIZE OF HETEROGENEITIES

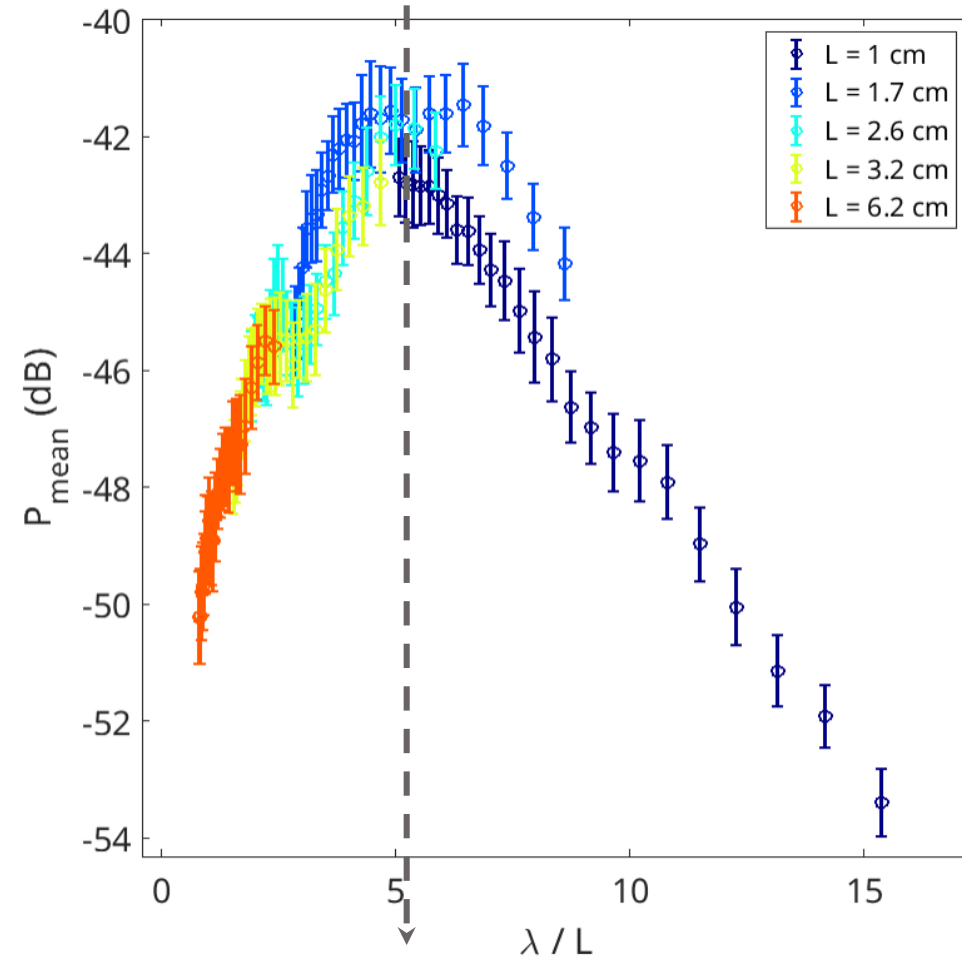
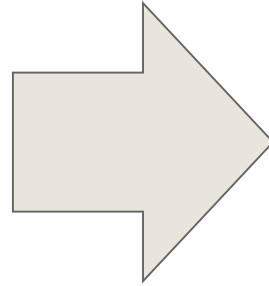
Bi-Modal distribution of permittivity

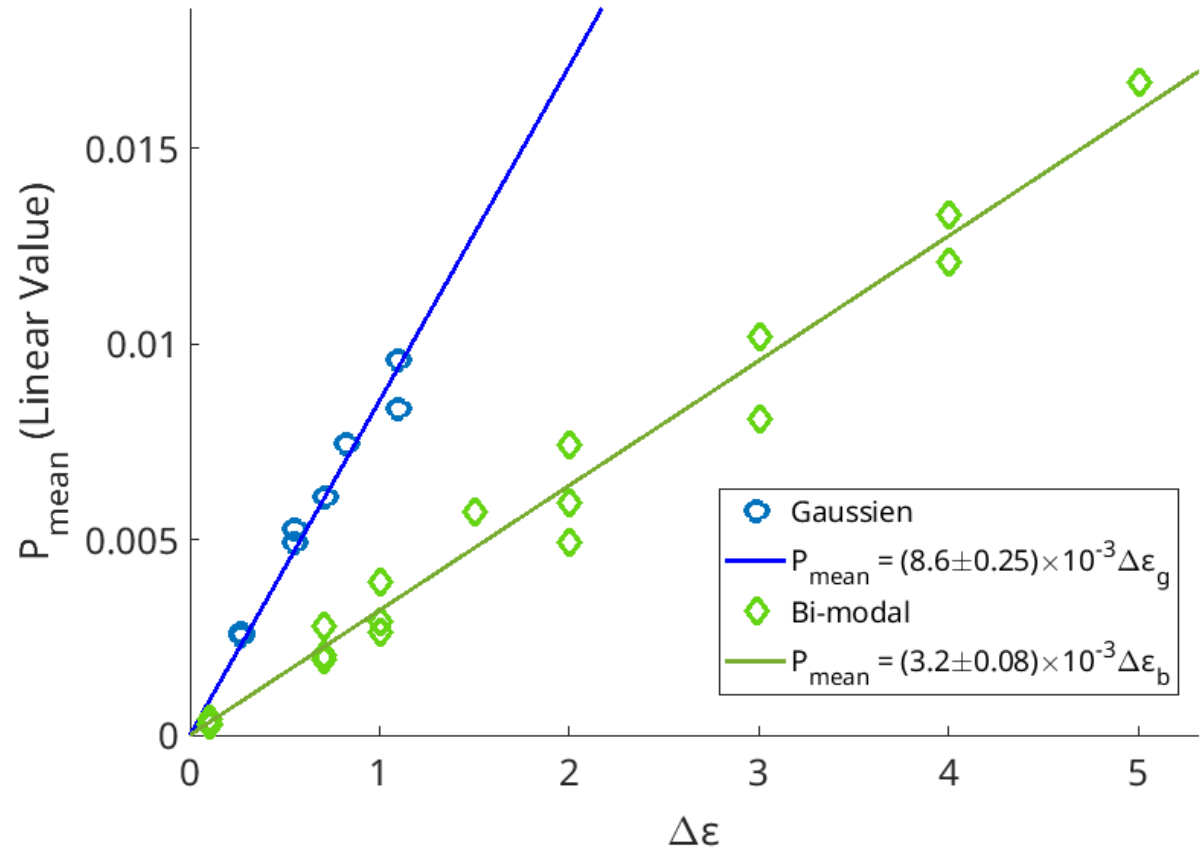
$\epsilon = 6$

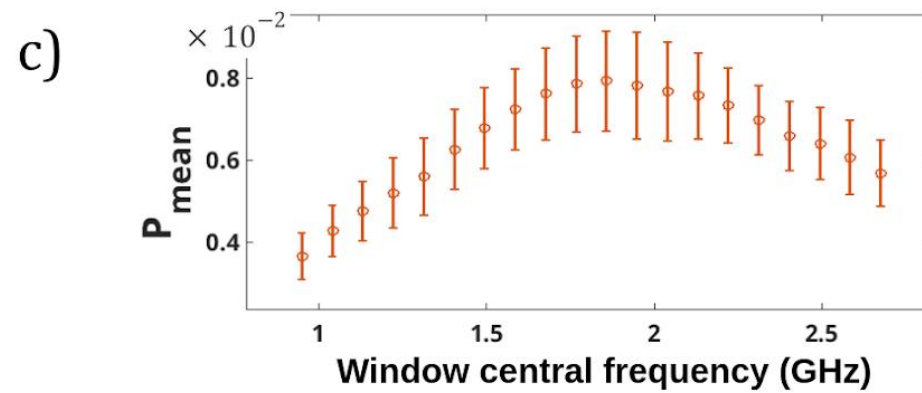
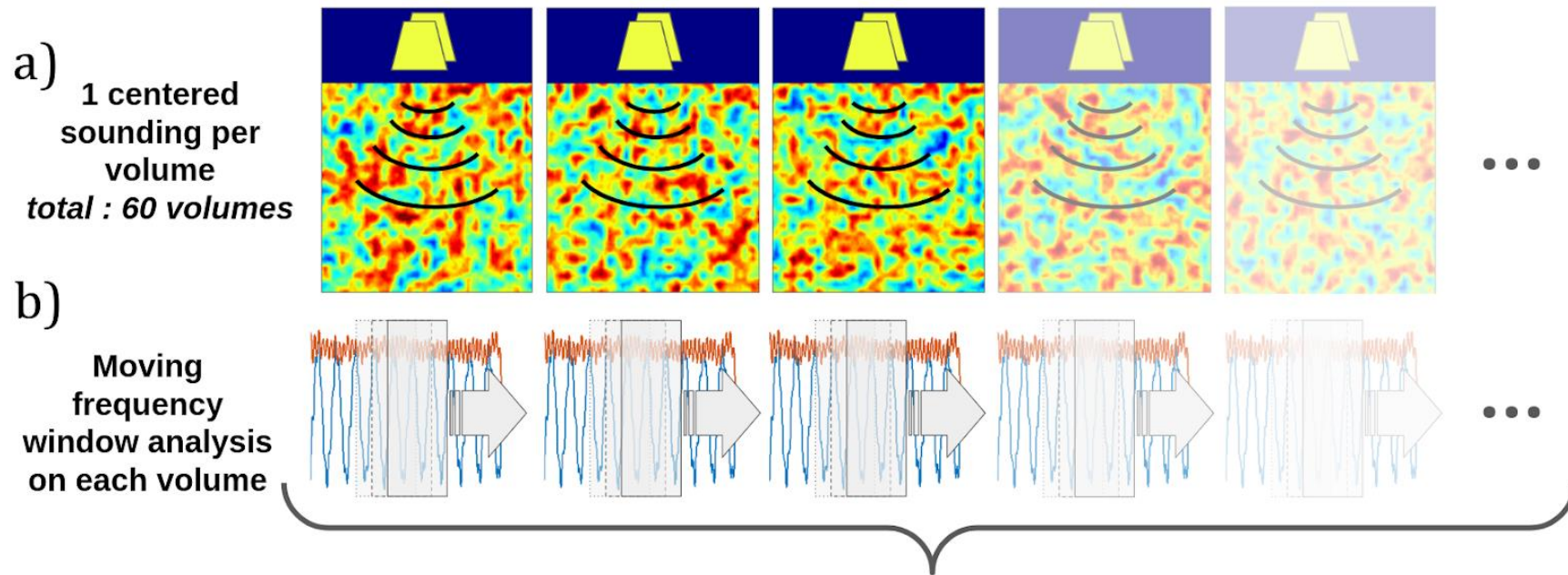
$\epsilon = 4$



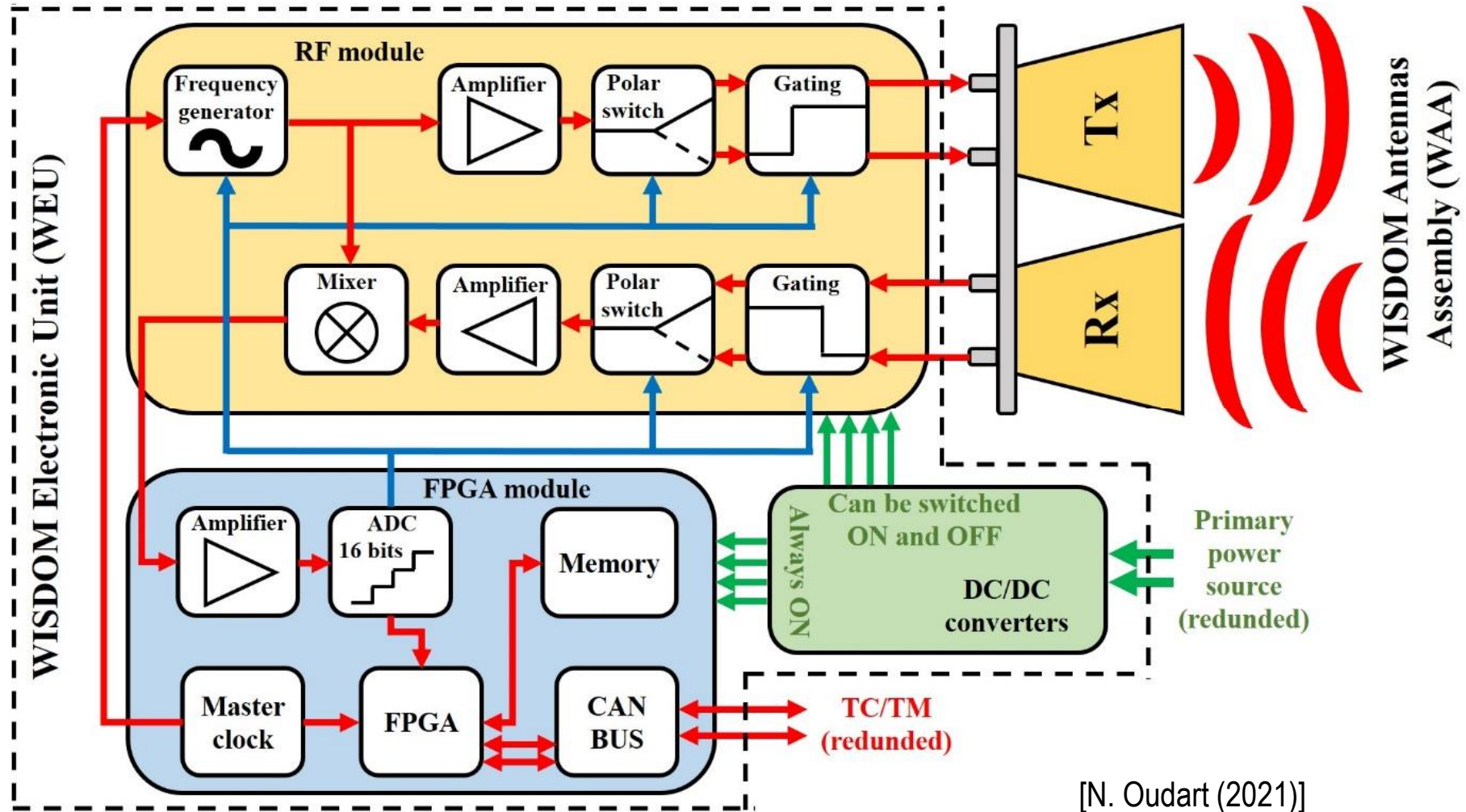
(L = 2.6 cm)







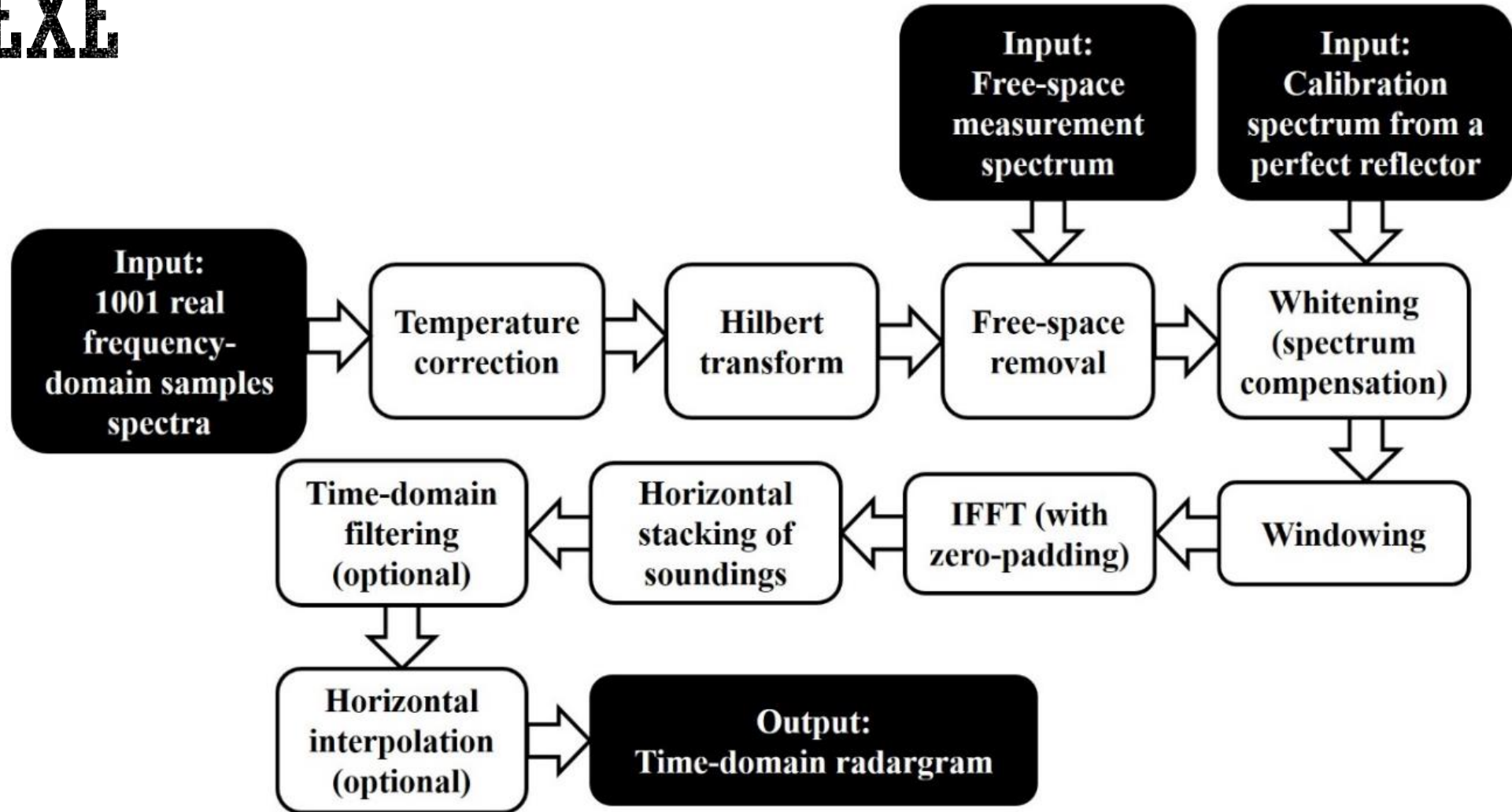
ANNEXE



[N. Oudart (2021)]



ANNEXE



[N. Oudart (2021)]



ANNEXE

	WISDOM
Radar design	Stepped Frequency Continuous Wave
Working frequencies	0.5 – 3 GHz
Antennas design	Vivaldi antennas
Free-space resolution (worst case)	6 cm
Expected penetration depth (material dependent)	3 – 10 m
Dynamic range	~84 dB of effective range
Power consumption	12.5 W
Mass	1.36 kg

